

First-Order Flow Equations for Extremal Non-BPS Black Holes

Gabriel Lopes Cardoso

with Ceresole, Dall'Agata, Oberreuter, Perz, arXiv 0706.3373

3rd RTN Workshop, Valencia, 03.10.07

ARNOLD SOMMERFELD
CENTER FOR THEORETICAL PHYSICS



Extremal Non-BPS Black Hole Solutions

Goal: construct extremal, **non-BPS** black hole solutions in $N = 2, D = 4$ sugra theories at 2-derivative level, by

- solving **first-order** flow equations;
- use **5D/4D-connection** for extremal BHS to achieve this
Gaiotto+Strominger+Yin, hep-th/0503217
 - ▶ single-center, static solutions to $N = 2, D = 4$
 - ▶ dyonic $(p^0, p^A; q_0, q_A), \quad A = 1, \dots, n$
 - ▶ supported by complex scalar fields z^A
 - ▶ spherical symmetry: metric factor $U = U(r)$
 $z^A = z^A(r)$

Supersymmetric Case

BPS case: solutions obtained by solving first-order differential flow equations associated with $\delta \text{Fermion} = 0$:

Ferrara+Gibbons+Kallosh+Strominger 1996

$$\begin{aligned}\partial_\tau e^{-U} &= |Z_4| , \\ \partial_\tau z^A &= -2 e^U g^{A\bar{B}} \partial_{\bar{B}} |Z_4| ,\end{aligned}$$

where $\tau = 1/r$, and $Z_4 = Z_4(p, q, z^A)$ central charge.

Solutions expressed in terms of harmonic functions ($H^0, H^A; H_0, H_A$).

Flow: spatial infinity $r = \infty, z^A(\infty)$; horizon $r = 0, z^A(0)$.

Under **good conditions**, flow in moduli space from $z^A(\infty)$ to $z^A(0)$ evolves smoothly to a **fixed-point** at the horizon with $Z_4 \neq 0$:

attractor point $\partial_{\bar{B}} |Z_4| = 0$, entropy $S_{\text{macro}} = \pi |Z_4|_{\text{hor}}^2$

Non-Supersymmetric Case

There exist **non-BPS** black hole solutions that can be constructed by solving **first-order** flow equations.

Ceresole+Dall'Agata, hep-th/0702088

Associated flow equations:

$$\begin{aligned}\partial_\tau e^{-U} &= W_4, \\ \partial_\tau z^A &= -2 e^U g^{A\bar{B}} \partial_{\bar{B}} W_4,\end{aligned}$$

with W_4 **real**, $W_4 \neq |Z_4|$.

Under **good conditions**, flow has again a **fixed-point** at horizon with $W_4 \neq 0$:

attractor point $\partial_{\bar{B}} W_4 = 0$, entropy $S_{\text{macro}} = \pi W_{4\text{hor}}^2$

Effective Black Hole Potential

When do such first-order flow equations exist?

Consider the effective **black hole potential**

$$F^2 \longrightarrow V_{\text{BH}} = Q^T \mathcal{M}(z, \bar{z}) Q \quad , \quad Q^T = (p, q) .$$

Whenever

$$V_{\text{BH}} = W_4^2 + 4 g^{A\bar{B}} \partial_A W_4 \partial_{\bar{B}} W_4 ,$$

can rewrite 2-derivative Lagrangian (dim. reduced on spher. symm. background) as

$$\int L_4 = \sum \int (\text{first - order flow equations})^2 + \int \text{T.D.}$$

Solution to first-order eqs satisfies EOM.

See also [Andrianopoli+D'Auria+Orazi, Trigiante 0706.0712](#)

Effective Black Hole Potential

Rewriting of V_{BH} is **not** unique:

- if in terms of $Z_4 \longrightarrow$ BPS
- if in terms of $W_4 \neq |Z_4| \longrightarrow$ non-BPS

Here: construct W_4 by using the **5D/4D-connection** for extremal black holes. Gaiotto+Strominger+Yin, hep-th/0503217

Rotating extremal BHS in $D = 5$ in **Taub-NUT space** \leftrightarrow static extremal BHS in $D = 4$.

In the vicinity of NUT charge, spacetime looks five-dimensional, whereas far away from it spacetime looks four-dimensional.

VSG ($N = 2, D = 5$) simpler to deal with.

Related work by Ceresole+Ferrara+Marrani, 0707.0964

Static Extremal BHS in Taub-NUT

First: **static** extremal 5D BHS in Taub-NUT space.

Electric charges q_A , Taub-NUT charge $p^0 > 0$

Effective **black hole potential** in $D = 5$:

$$G_{AB} F^A F^B \longrightarrow V_{\text{BH}} = q_A G^{AB} q_B$$

Suppose G^{-1} possesses an **invariance** group with elements R^A_B , i.e.

$$R^T G^{-1} R = G^{-1} , \quad R = \text{const} , \text{ real}$$

Then

$$V_{\text{BH}} = \frac{2}{3} W_5^2 + G^{AB} \partial_A W_5 \partial_B W_5$$

with

$$W_5 = Q_A X^A , \quad Q_A = q_B R^B_A$$

- if $R^A_B = \delta^A_B \longrightarrow W_5 = Z_5$, BPS black hole in 5D
- if $R^A_B \neq \delta^A_B \longrightarrow W_5 \neq Z_5$, non-BPS black hole in 5D

Static Extremal BHS in Taub-NUT

Ex: G_{AB} diagonal, then $R = \text{diag}(\pm 1, \mp 1, \dots, \pm 1)$

Can now rewrite

$$\int L_5 = \sum \int (\text{first - order flow equations})^2 + \int \text{T.D.}$$

→ solved in terms of harmonic functions ($\tau = 1/r$)

$$H_A = h_A + |Q_A| \tau, \quad h_A > 0$$

Reducing to 4D over Taub-NUT circle yields first-order flow eqs. for a static extremal dyonic BH with charges (p^0, q_A) , based on

$$W_4 = \frac{1}{8} e^{K/2} \left(e^{-K} p^0 + 4 |Q_A (z^A - \bar{z}^A)| \right)$$

When $R^A_B \neq \delta^A_B$, $W_4 \neq |Z_4|$ → **non-BPS** BH in 4D

4D BH solution is constructed out of harmonic functions ($\tau = 1/r$)

$$H^0 = h^0 + p^0 \tau, \quad H_A = h_A + |Q_A| \tau$$

Rotating Extremal BHS in Taub-NUT

Now: consider **rotating** extremal 5D BHS in Taub-NUT space.

Stationary part $\omega = \omega_5 d\psi$ of line element along Taub-NUT circle ψ .

Chern-Simons term $C_{ABC} F^A \wedge F^B \wedge F^C$ in L_5 now contributes. Its **first-order rewriting** exists if R^A_B satisfies, in addition, the relation

$$2G_{AB} R^B_C X^C = C_{ABC} R^B_D X^D R^C_E X^E \quad (*)$$

Then, denoting $s = \text{sgn}(W_5)$, $W_5 = Q_A X^A$,

- $s = 1$, $\omega_5 = H_0 = h_0 + q_0 \tau$, harmonic
- $s = -1$, $\omega_5 = c/H^0$, $c = \text{const}$

Ex: take $R^A_B = \delta^A_B$ (satisfies $(*)$) and $s = -1$, and reduce over Taub-NUT direction ψ . This yields a 4D flow based on

$$W_4 = e^{K/2} \left| \frac{1}{6} \rho^0 C_{ABC} z^A z^B \bar{z}^C - q_A z^A \right| \neq |Z_4|$$

Since $\omega_5 \neq 0$ along flow, have non-trivial **axions** $\text{Re}z^A$ along the flow.

- 1 Using $5D/4D$ -connection, constructed extremal non-BPS BHS with charges (p^0, q_A) . Apply duality transformations to obtain other non-BPS BHS, for instance (q_0, p^A) .
- 2 An example of an extremal non-BPS BH **not** captured by the above construction:
 $4D (p^0, q_0) \longleftrightarrow 5D (p^0, J)$ Kerr BH
- 3 Multiple-center solutions? Split attractor flow for non-BPS black holes?