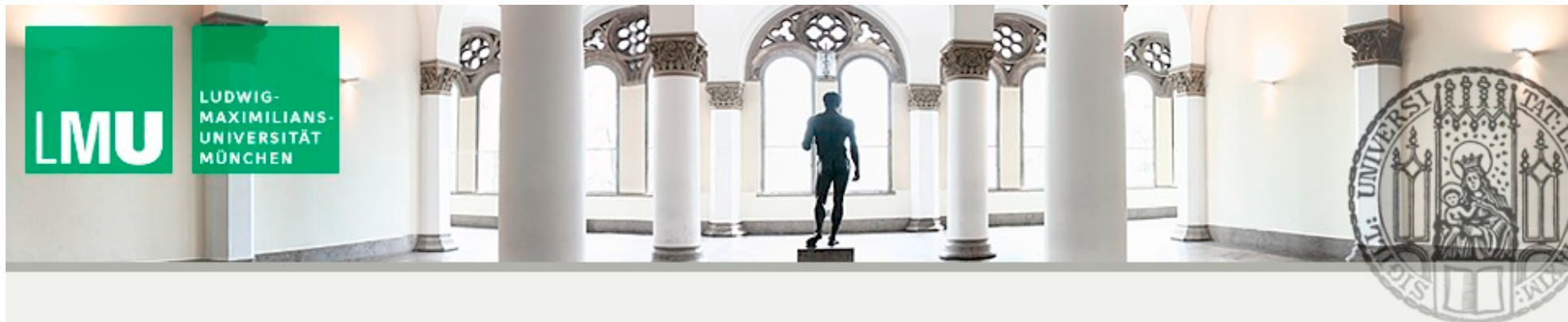


The Landscape of String theory

Dieter Lüst, LMU (ASC) and MPI München



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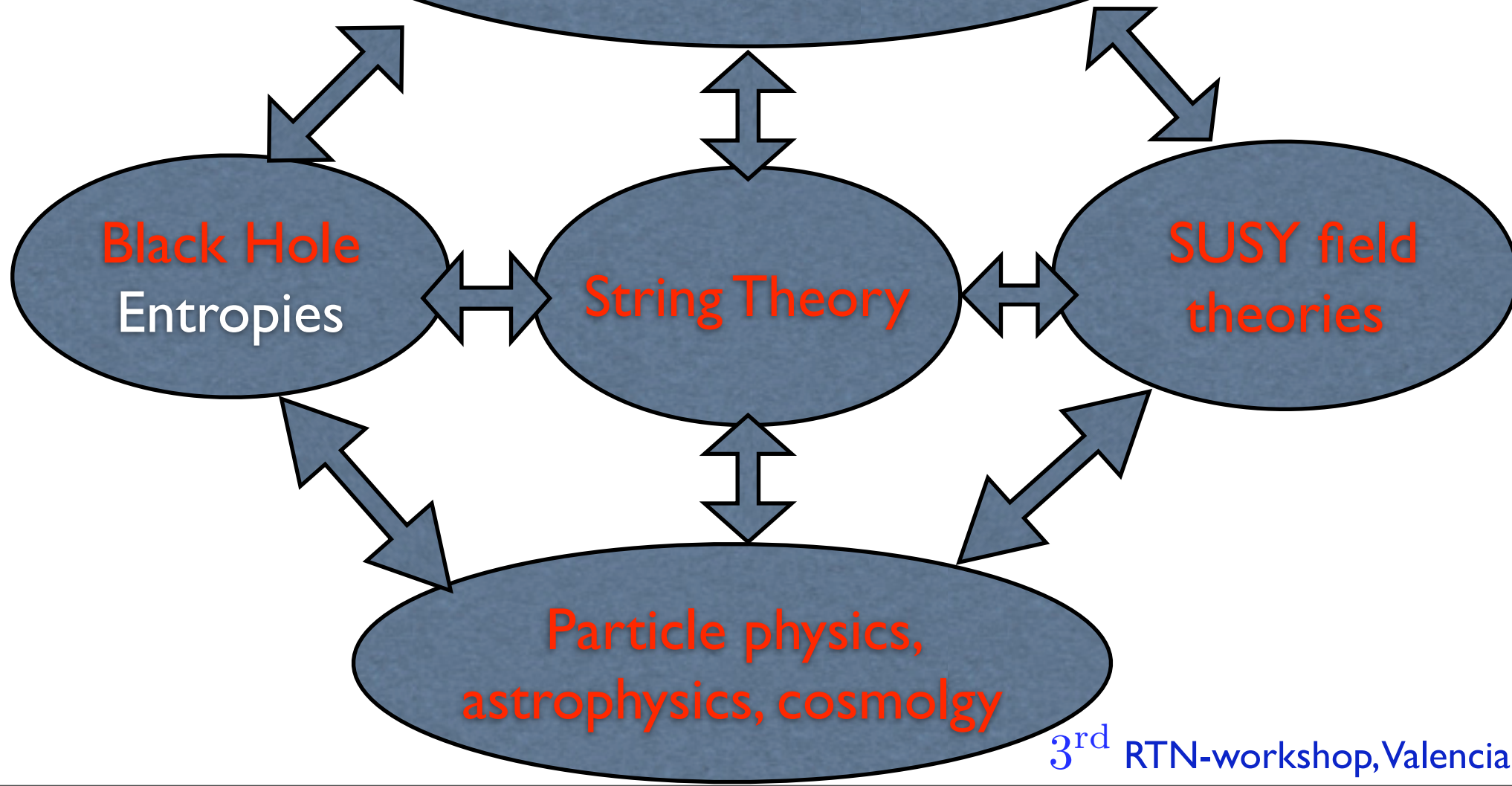
Dieter Lüst, LMU (ASC) and MPI München

in collaboration with

Riccardo Appreda, Ralph Blumenhagen, Gabriel L. Cardoso, Mirjam Cvetič, Johanna Erdmenger, Florian Gmeiner, Viviane Grass, Michael Haack, Daniel Krefl, Gabriele Honecker, Costas Kounnas, Jan Perz, Marios Petropoulos, Susanne Reffert, Robert Richter, Christoph Siegel, Maren Stein, Stephan Stieberger, Antoine van Proeyen, Dimitri Tsimpis, Timo Weigand and Marco Zagermann

I) Introduction

Geometry: Calabi-Yau spaces, mirror symmetry, generalized spaces, D-branes, K-theory, ...



String theory:



Describes the interactions and the spectrum of D -dimensional extended (closed & open) strings.

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- Intermediate energies ($E \sim \mathcal{O}(\text{TeV})$, $D = 4(6?)$) :

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Perturbative (supersymmetric?) **Standard Model**
(by compactification) with **3 quark families**.
- Low energies ($E \sim \mathcal{O}(\text{GeV})$, $D = 4(5?)$)
Can string theory provide some non-perturbative informations for QCD?

Count the number of consistent string solutions



Vast landscape with $N_{sol} = 10^{500-1500}$ discrete vacua!

(Lerche, Lüst, Schellekens (1986), Douglas (2003))



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- Explore all mathematically consistent possibilities:
top down approach (quite hard), string statistics
(perhaps some anthropic point of view is necessary?)

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Two strategies to find something interesting:

- Explore all mathematically consistent possibilities: **top down approach** (quite hard), string statistics (perhaps some anthropic point of view is necessary?)
- Do not look randomly - look for green (promising) spots in the landscape ➔ model building, **bottom up approach**.

- Heterotic string compactifications
- Type II orientifolds models
 - Intersecting brane models and their statistics
 - D-instantons: non-perturbative couplings
- The string landscape: fluxes and branes
 - AdS_2 flux vacua and black holes
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4D chiral matter fields: $N_F = c_3(V)$





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Resume: heterotic strings:





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- ☹ Moduli stabilization (bundle moduli!) with H-flux is difficult.

(Strominger (1985), Becker, Becker, Dasguta, Green (2003); Curio, Cardoso, Dall'Agata, Lüst, Krause (2003/04/05); Braun, He, Ovrut (2006))

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- Can be combined with background fluxes \Rightarrow **moduli stabilization (GKP) and dS-vacua (KKLT)**

(Review: Blumenhagen, Körs, Lüst, Stieberger, hep-th/0610327)





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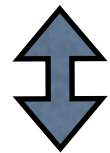
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- (diophantic equations with finite no. of solutions!)



Geometrical, large radius regime:

IIA: special lagrangian submanifolds: D6 on 3-cycles **at angles**



Mirror symmetry (SYZ)

IIB: points, (complex lines), divisors, (CY) **with gauge bundles:**

D3

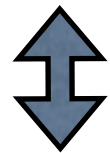
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D7

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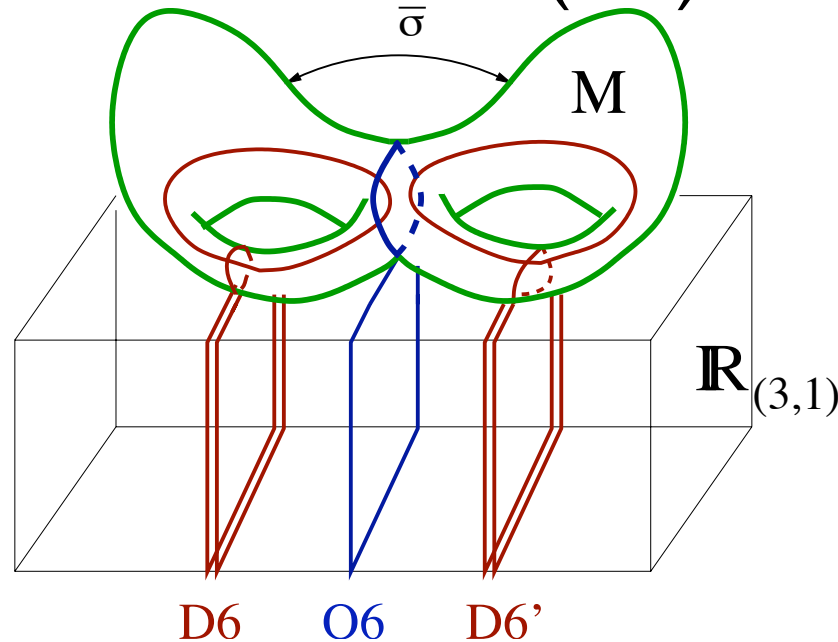
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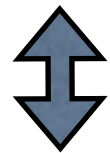
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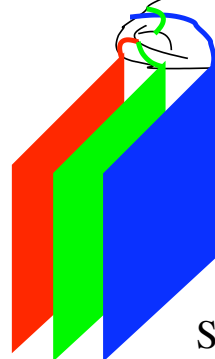
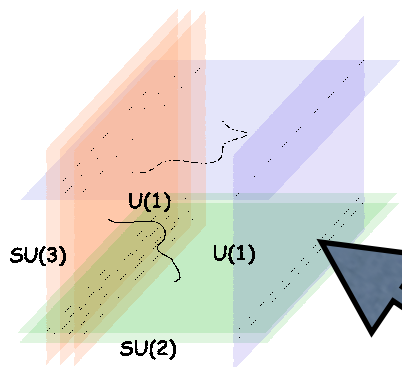
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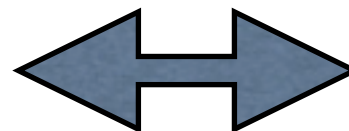
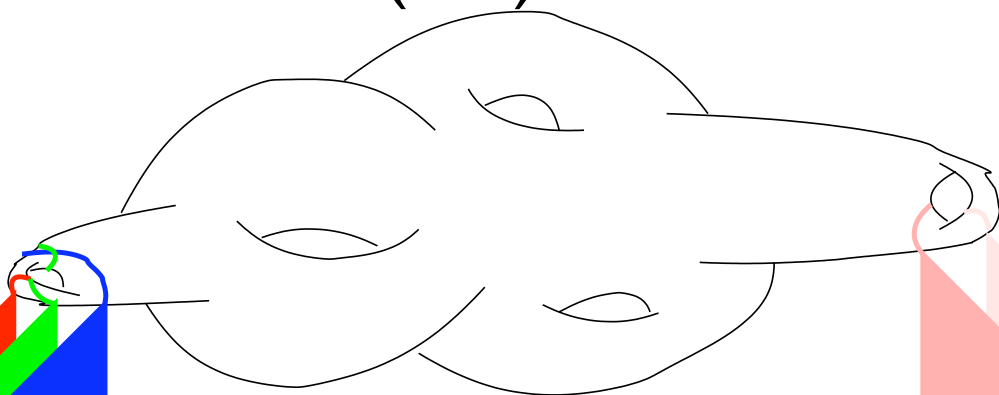
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SM

Soft SUSY breaking



HS



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(Blumenhagen, Gmeiner, Honecker, Lüst, Stein, Weigand, [hep-th/0411173](#), [hep-th/0510170](#), [hep-th/0703011](#);
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Example: $\mathcal{M}_6 = T^6 / (Z_2 \times Z_2)$ IIA orientifold:

Systematic computer search (NP complete problem):

Look for solutions of a set of diophantic equations:

There exist about $1.66 \cdot 10^8$ susy D-brane models on this orbifold (with restricted complex structure)!

(Finiteness of models was recently proven by D.T.)



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Restriction	Factor
gauge factor $U(3)$	0.0816
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No symmetric representations	0.839
Massless $U(1)_Y$	0.423
Three generations of quarks	2.92×10^{-5}
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**Only one in a billion models gives
rise to a MSSM like vacuum!**



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- **Explicit D-brane constructions:**
there exist many models that come close to the MSSM.
Problem of exotic particles!

(Chen, Li, Mayes, Nanopoulos, hep-th/0703280; Chen, Li, Nanopoulos, hep-th/0604107; Blumenhagen, Plauschinn, hep-th/0604033; Bailin, Love, hep-th/0603172; Blumenhagen, Cvetic, Marchesano, Shiu, hep-th/0502095; Marchesano, Shiu, hep-th/0409132; Honecker, Ott, hep-th/0407181;

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- **Unbroken space-time supersymmetry.**

Take into account non-perturbative instanton corrections to the effective action!





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- Can generate new matter couplings (Majorana masses, Yukawa couplings) → see in a moment.





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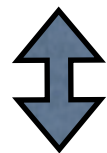


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IIA: special lagrangian submanifolds: E2 on 3-cycles



Mirror symmetry (SYZ)

II B: points, (complex lines), divisors, (CY)

E(-1) (E1) E3 (E5)

Open string instanton CFT:

(R. Blumenhagen, M. Cvetič, T. Weigand)



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- Finally, one has to integrate over all bosonic and fermionic zero modes localised on $E(p)$ -branes
- Not all instantons can contribute to the effective action:
 - F-terms: $E(p)$ -brane is half-BPS:
need two fermionic zero modes θ_i
 - D-terms: $E(p)$ -brane breaks all 4 supersymmetries:
need four fermionic zero modes $\theta_i, \bar{\theta}_i$



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W_{np} must contain charged matter fields:

$$W_{np} \sim \prod \Phi_a e^{-S_E}$$

(U(1) selection rules will follow from fermionic zero modes.)

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Compute the I-instanton Veneziano-Yankielowicz/Affleck-Dine-Seiberg (VY/ADS) superpotential from string theory.

$$\begin{aligned}
 W_{\text{n.p.}} &= \frac{1}{\det[\tilde{\Phi}_{ab}\Phi_{ab}]} \exp - \left(f_{\text{a,tree}} + f_{\text{a,1-loop}}(M) \right) \\
 &= \frac{1}{\det[\tilde{\Phi}_{ab}\Phi_{ab}]} C(T) \exp \left(-f_{\text{a,tree}}(S, U) \right) \quad SU(N_c), \quad N_f = N_c -
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From integration over bosonic zero modes, ADHM constraints.

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We also computed the ADS superpotential for
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Possible generalizations:

- $SU(N_c)$, $N_f \leq N_c - 1$, due to gaugino condensation?
- Non-rigid cycles ($b_1(\Xi) \neq 0$) ?

This corresponds to adjoint chiral fields in gauge theory, which must get a mass by fluxes.



LMU Non-perturbative Yukawa couplings:

(R. Blumenhagen, M. Cvetič, D. Lüst, R. Richter, T. Weigand, arXiv:0707.1871)



MAX-PLANCK-GESELLSCHAFT

SU(5) GUT intersecting D6-brane models:

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Consider two stack **a** and **b** of D6-branes:

$$G = U(5)_a \times U(1)_b = SU(5)_a \times U(1)_a \times U(1)_b$$

Open strings:

sector	number	$U(5)_a \times U(1)_b$ reps.	$U(1)_X$
(a', a)	$3 + (1, 1)$	$\mathbf{10}_{(2,0)}$	$\frac{1}{2}$
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Abelian symmetries: $U(1)_a \times U(1)_b$

One anomalous linear combination: global symmetry

One anomaly free linear combination $U(1)_X$

Two different cases:

$U(1)_X$ massive: Georgi-Glashow model

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Perturb. allowed couplings (e.g. u,c,t-quarks): $\langle 10_{(2,0)} \bar{5}_{(-1,1)} \bar{5}_{(-1,-1)}^H \rangle$,

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These can be generated by E2-instantons:

The instanton has to wrap a rigid 3-cycle Ξ

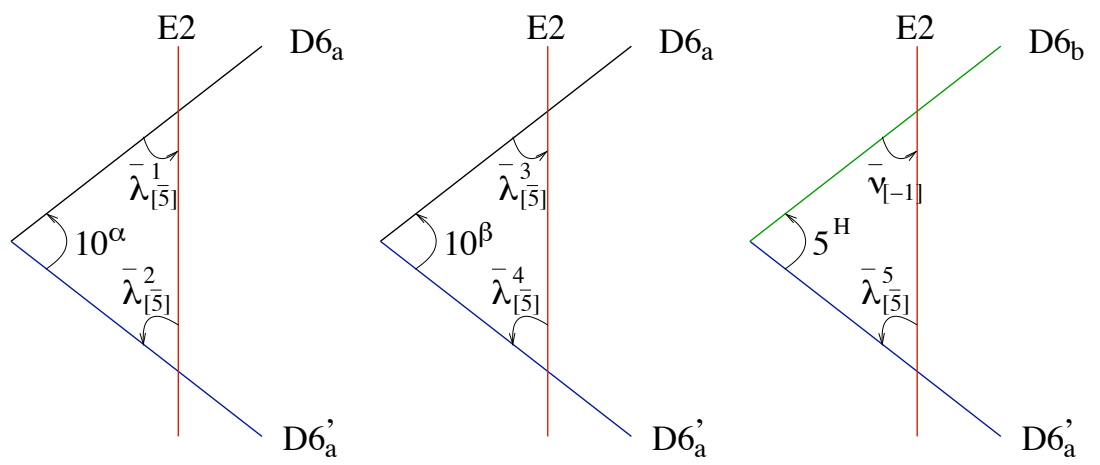
invariant under the orientifold projection $\Omega\bar{\sigma}$

and carrying gauge group $O(1)$

$U(1)_X$ charge of Yukawa coupling:
 one needs 6 fermionic zero modes:

Instanton intersection numbers:

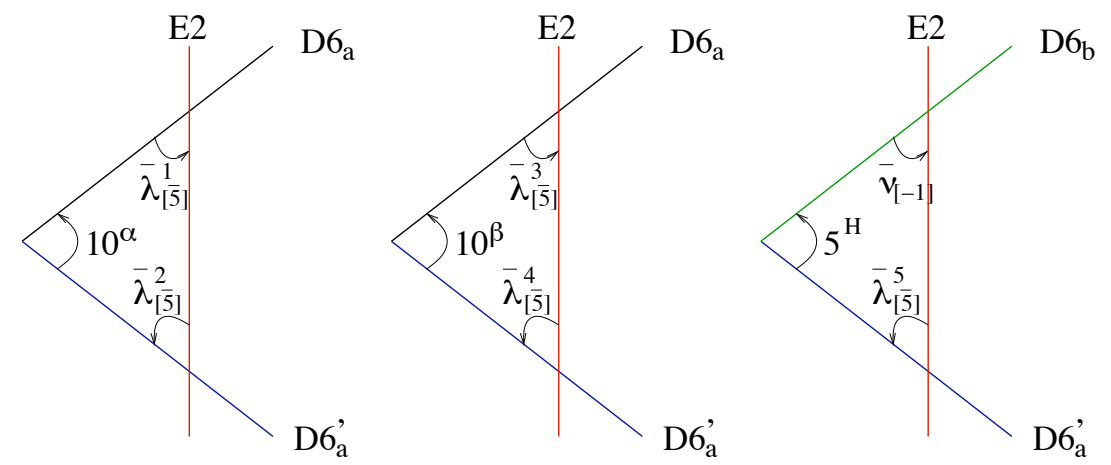
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Instanton action: $\exp(-S_{inst}) = \exp\left(C_\alpha^{10} 10_{[ij]}^\alpha \bar{\lambda}^i \bar{\lambda}^j + C^5 5_m \bar{\lambda}^m \bar{\nu}\right)$

Yukawa coupling: $W_Y = Y_{\langle 10 10 5_H \rangle}^{\alpha\beta} \epsilon_{ijklm} 10_{ij}^\alpha 10_{kl}^\beta 5_m^H e^{-S_{E2}} e^{Z'}$

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E2-instantons also relevant for doublet-triplet splitting,
Majorana masses, masses for exotic states, ...

(Ibanez, Uranga, hep-th/0609213; Cvetič, Richer, Weigand, hep-th/0703028; Ibanez, Schellekens, Uranga, arXiv:0704.1079; Antusch, Ibanez, Macri, arXiv:07062132; Bianchi, Kiritsis, arXiv:0702015; Bianchi, Fucito, Morales, arXiv:0704.0784)

LMU String model building: What do we like to learn?



MAX-PLANCK-GESELLSCHAFT



- Where is the fundamental string scale?



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- Where is the scale of supersymmetry breaking?



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(How good is the chain between fundamental theory and the data?)

- Heterotic string compactifications
- Type II orientifolds models
 - Intersecting brane models and their statistics
 - D-instantons: non-perturbative couplings
- The string landscape: fluxes and branes
 - AdS_2 flux vacua and black holes
 - AdS_4 flux vacua and domain walls

- Heterotic string compactifications

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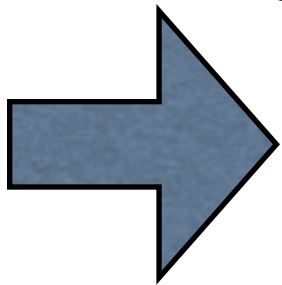
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(Weinberg, Bousso, Polchinski, Susskind, Linde, Schellekens, ...)

(Review: D. Lüst, arXiv:0707:2305)



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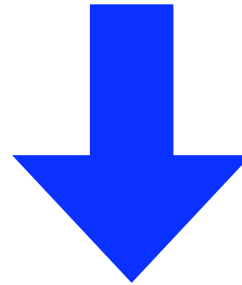
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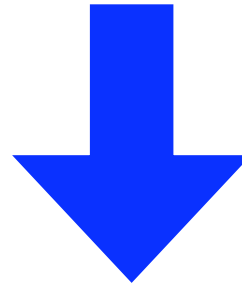
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Compactification

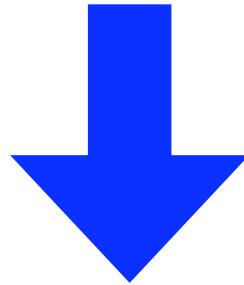
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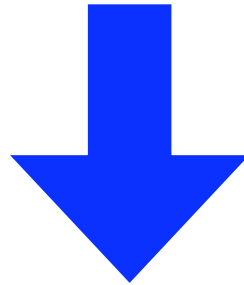


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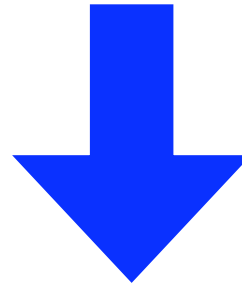
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How to make any prediction in string theory, i.e. how to determine the correct string vacuum state?

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However undetermined moduli result in uncalculable couplings (and new forces)!



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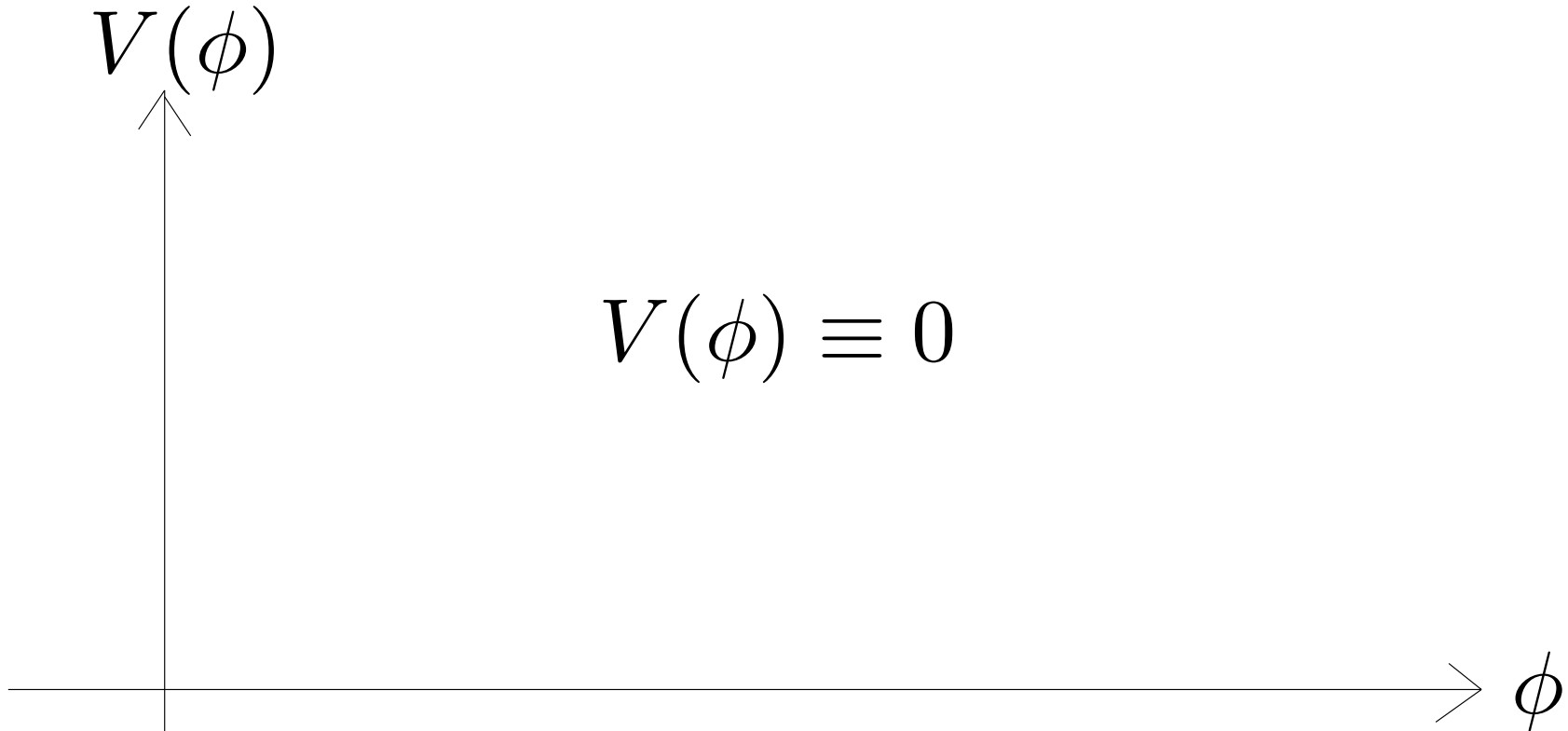
In this way one can obtain a discrete set of vacua with either

- negative cosmological constant, $\Lambda < 0$ (AdS vacua)
- zero cosmological constant, $\Lambda = 0$ (Minkowski vacua)
- positive cosmological constant, $\Lambda > 0$ (dS vacua)
- and various possibilities for gauge and matter fields

Effective moduli potential:



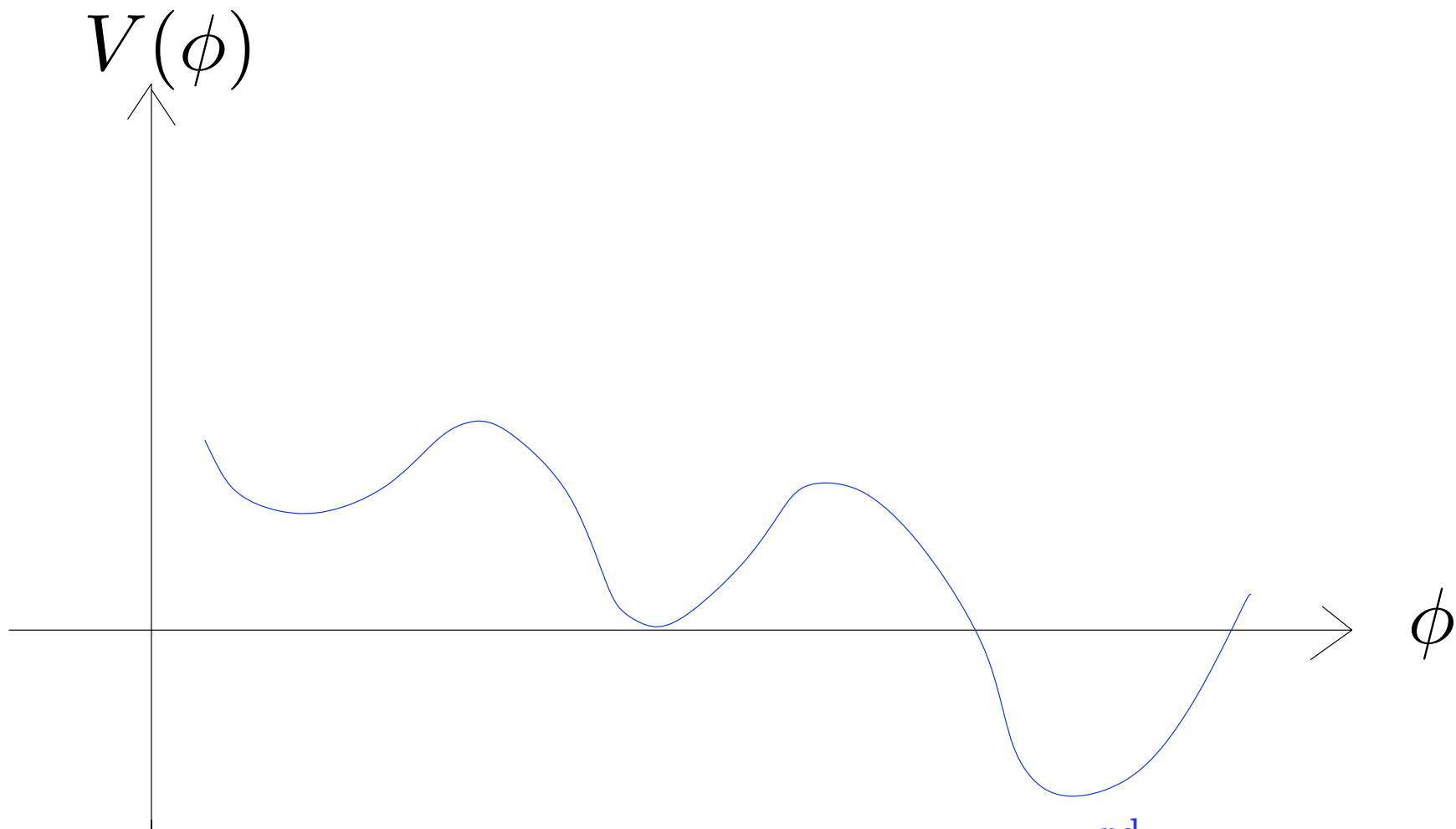
Flat moduli potential before turning on fluxes and non-perturbative effects:



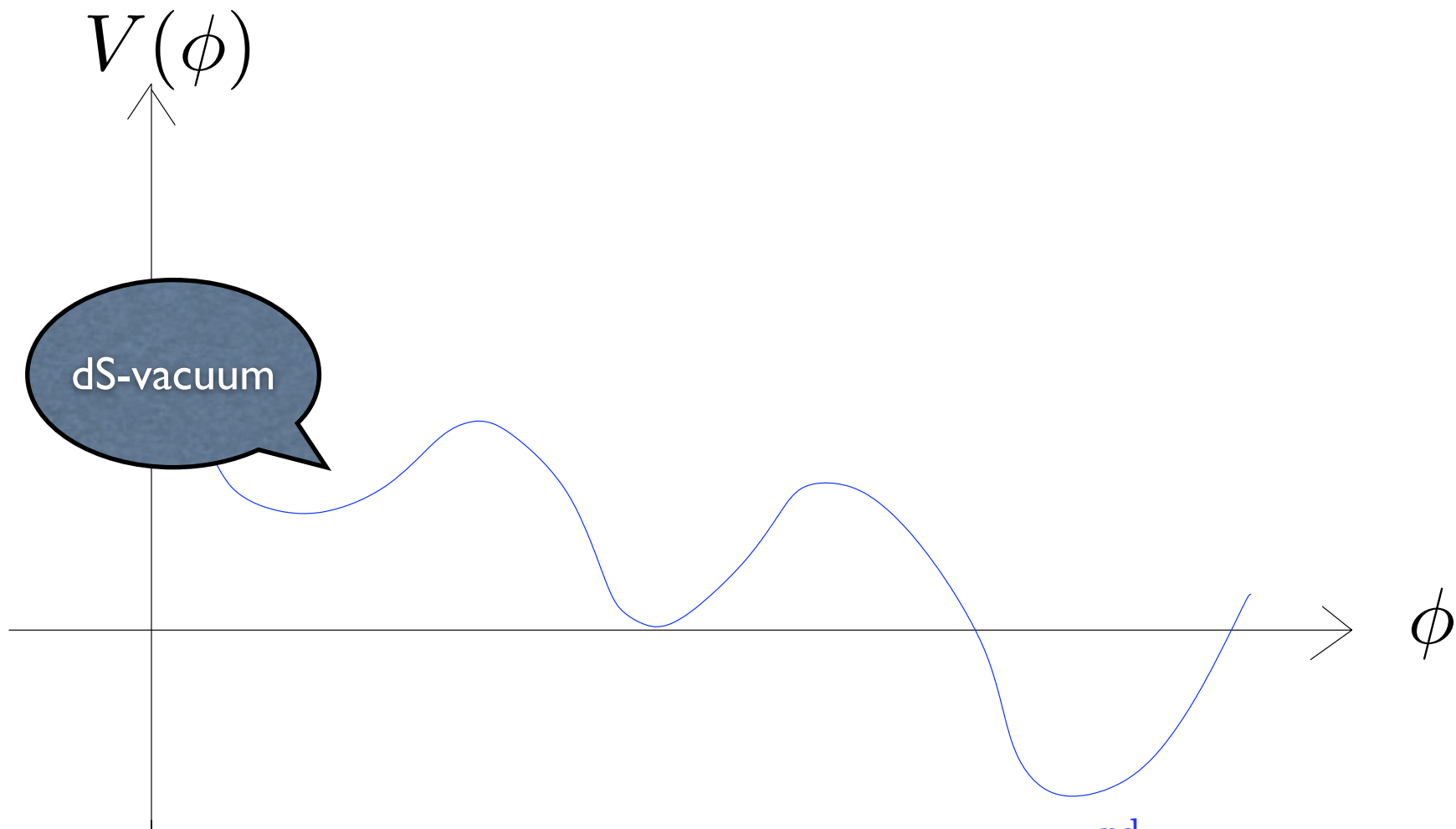
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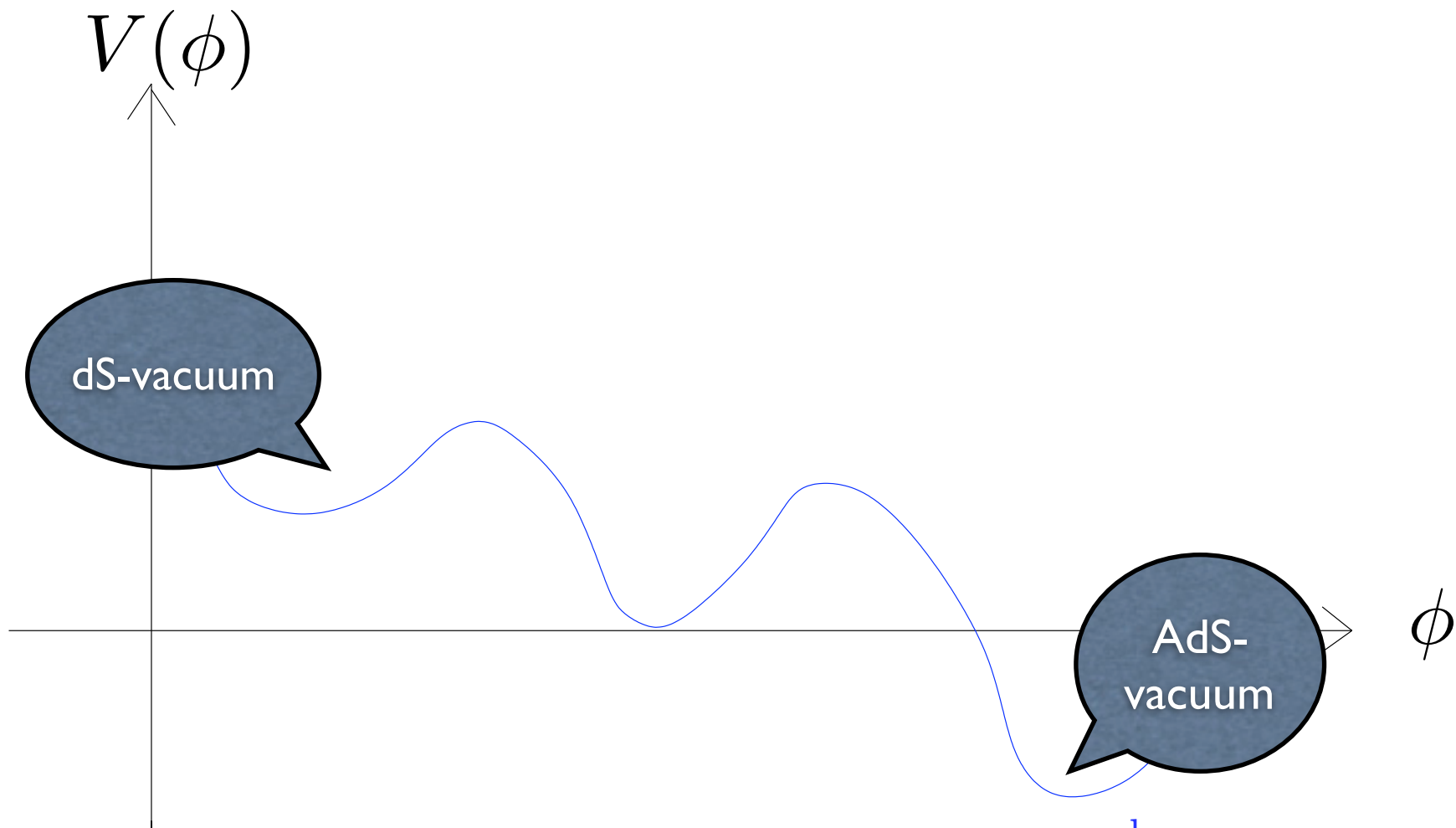
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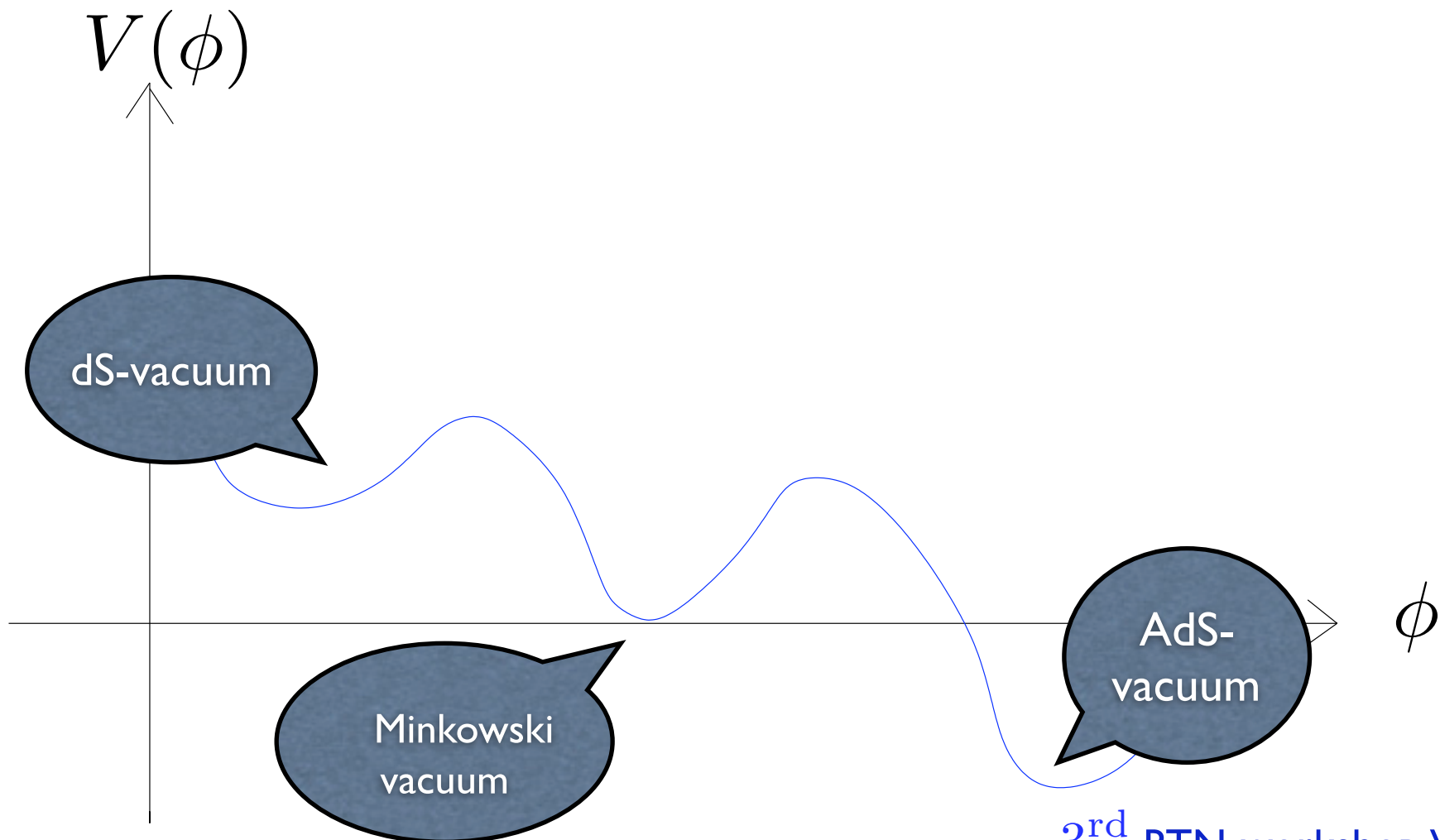
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$$\psi(\phi)$$

and see if $|\psi|^2$ is peaked, i.e. has maxima with good phenomenological properties.

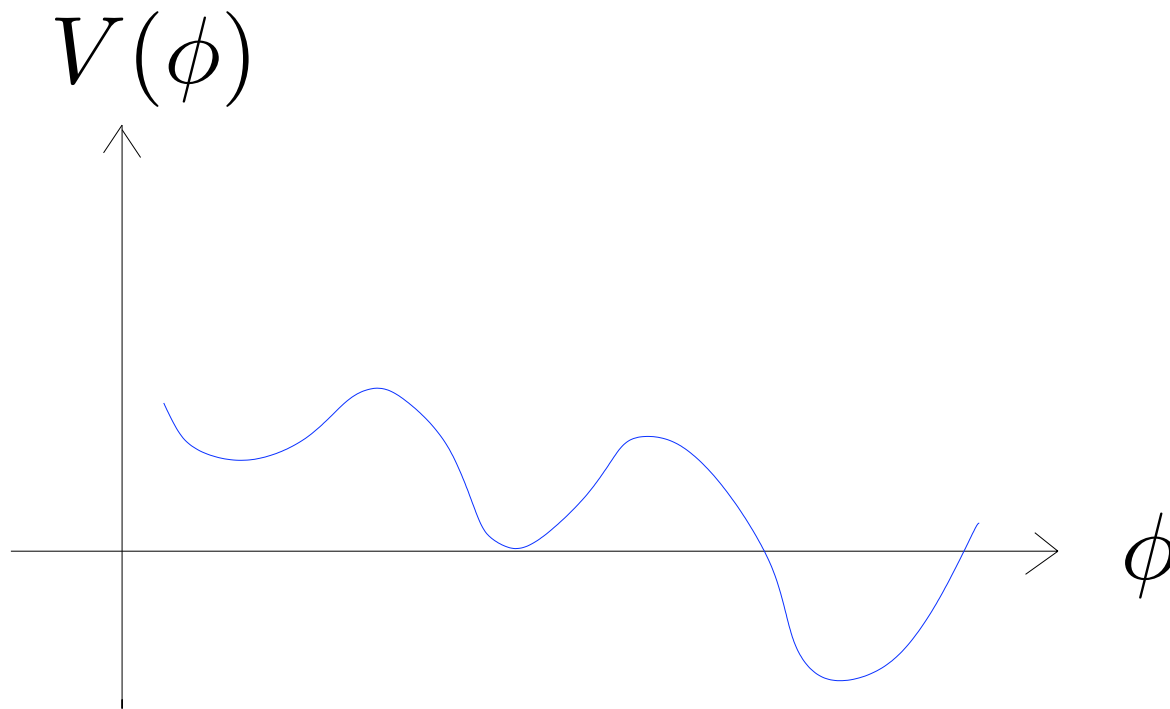


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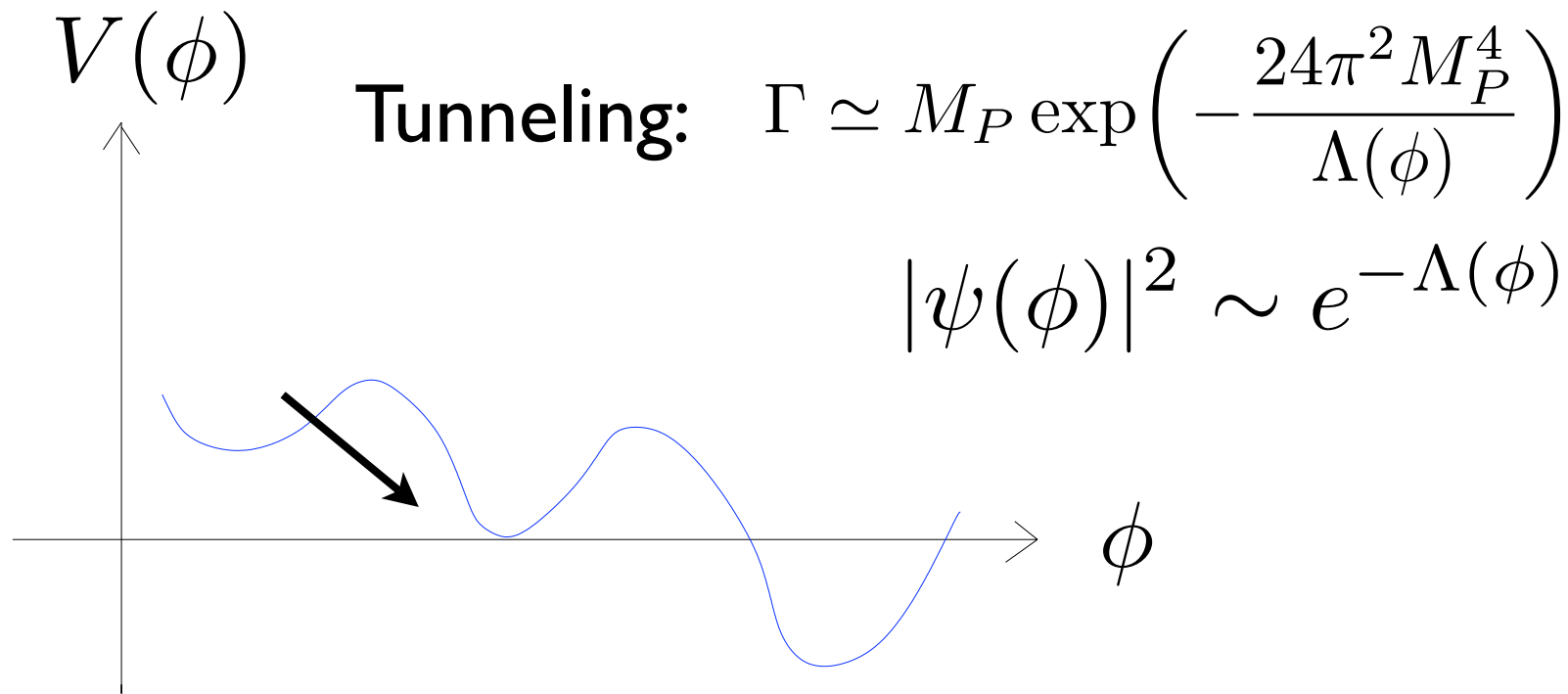
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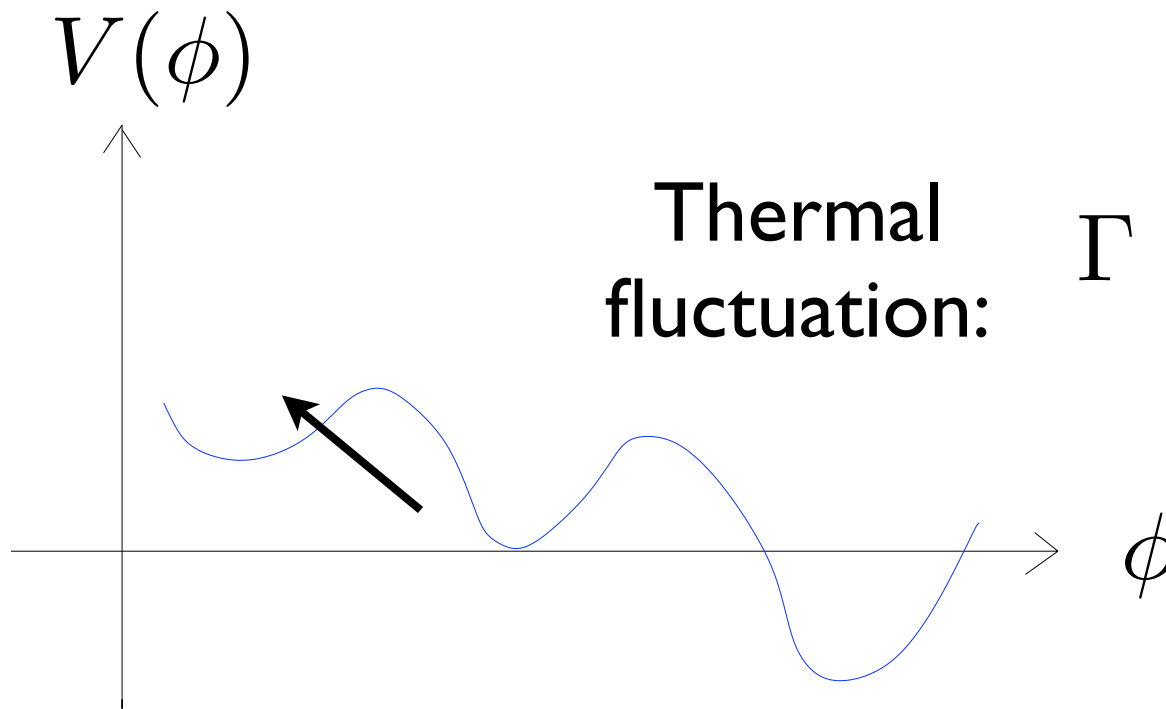
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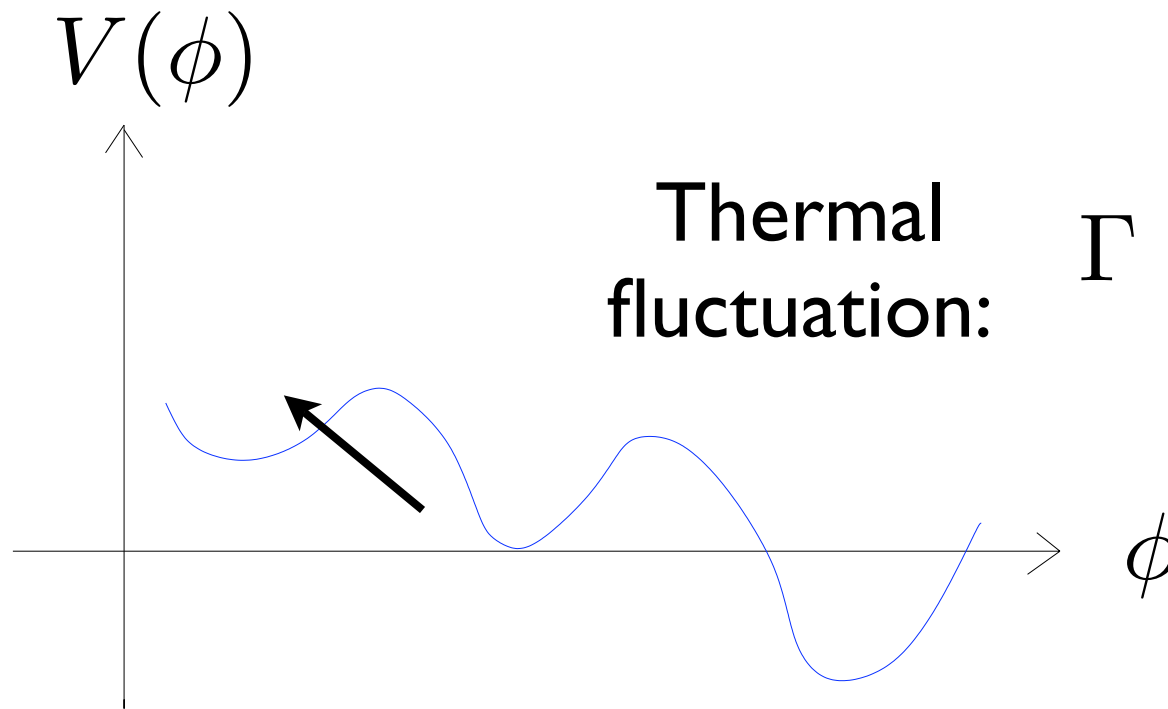
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(ii) „Stringy“ computation of the entropy of string vacua:

$$|\psi(\phi)|^2 = e^{\mathcal{S}_{\text{string}}(\phi)}$$

How to define an entropy for string vacua?





Similar problem in quantum gravity:

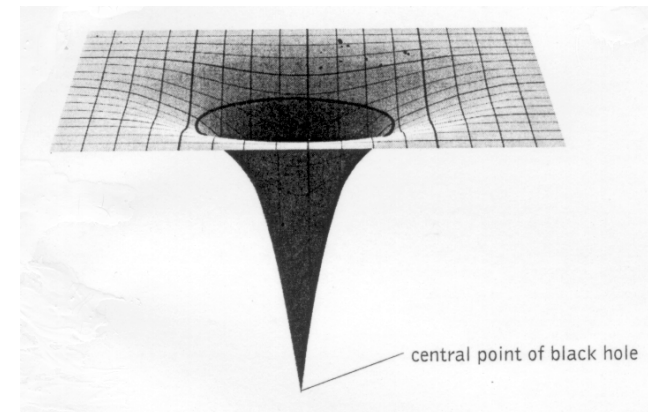
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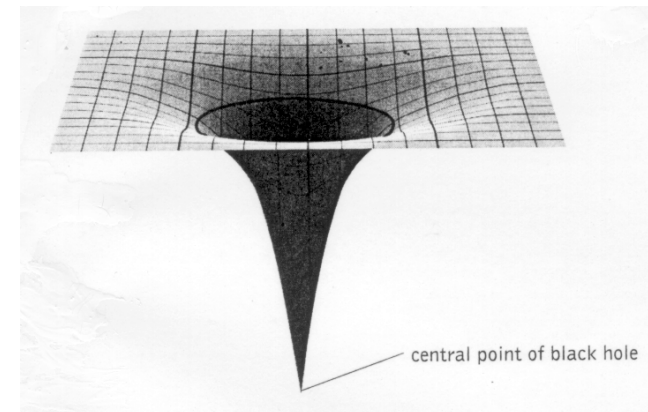
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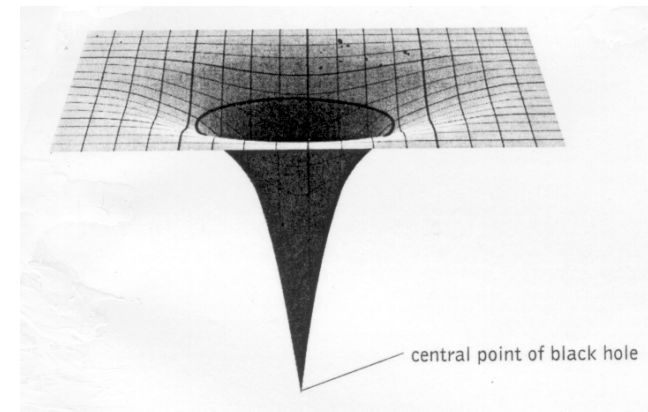


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Conjecture: (Ooguri, Vafa, Verline, hep-th/0502211)



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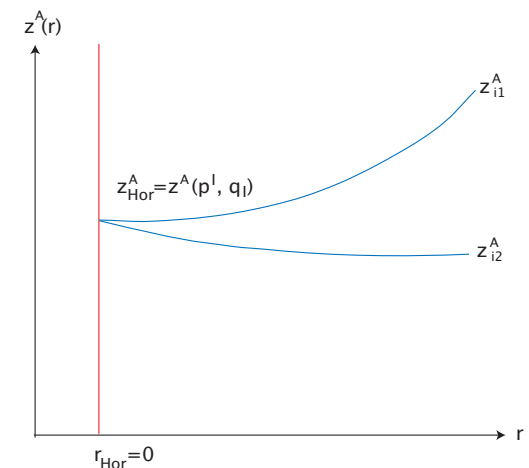
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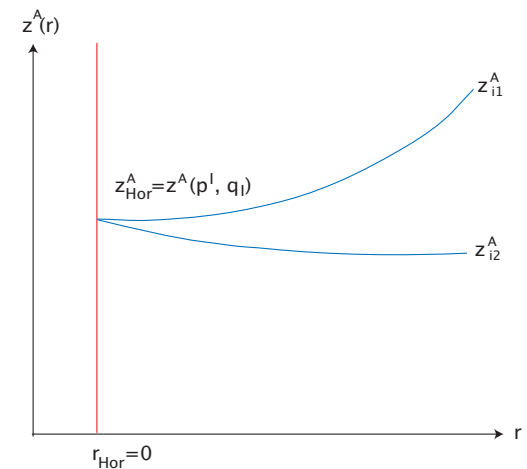
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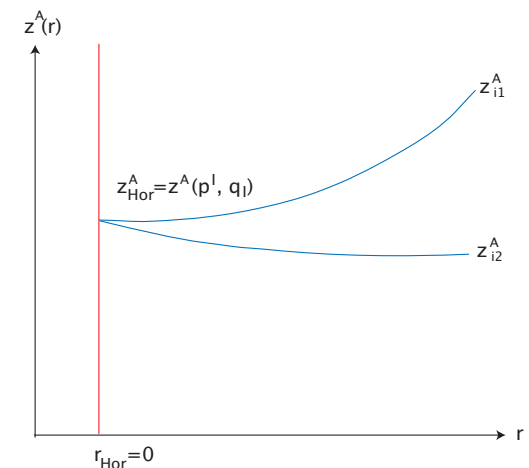
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(The black hole BPS conditions are very similar to the flux SUSY conditions, replacing the central charge Z by the superpotential W !)



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Near horizon geometry \Leftrightarrow Vacuum geometry

$$AdS_2 \times S^2 \times CY$$



(Cardoso, Lüst, Perz, [hep-th/0603211](#); related work: Gukov, Saraikin, Vafa, [hep-th/0509109](#);
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Following the entropic principle infrared free theories seem to be preferred!

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- Type II orientifolds models

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D-instantons: non-perturbative couplings

- The string landscape: fluxes and branes

AdS_2 flux vacua and black holes

AdS_4 flux vacua and domain walls

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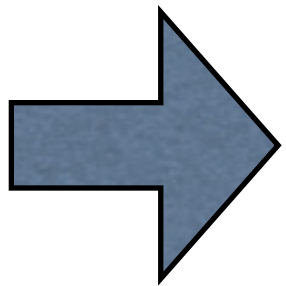
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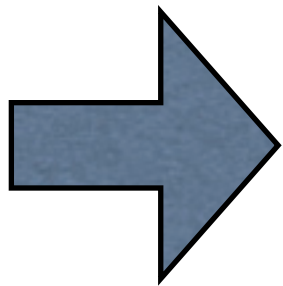
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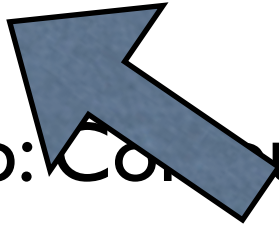
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AdS_4 flux vacua with all moduli fixed:





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without Kähler moduli (see later)



Tree level type IIA flux vacua:

(Curio, Klemm, Lüst, Theisen, hep-th/0012213; Derendinger, Kounnas, Petropoulos, Zwirner, hep-th/0411276; Villadoro, Zwirner, hep-th/0503169; De Wolfe, Giriyavets, Kachru, Taylor, hep-th/0505160; Camara, Font, Ibanez, hep-th/0506066; Villadoro, Zwirner, arXiv:0706.3049)



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0-form R-flux
(massive IIA
SUGRA)

2-form R-flux

4-form R-flux

6-form R-flux

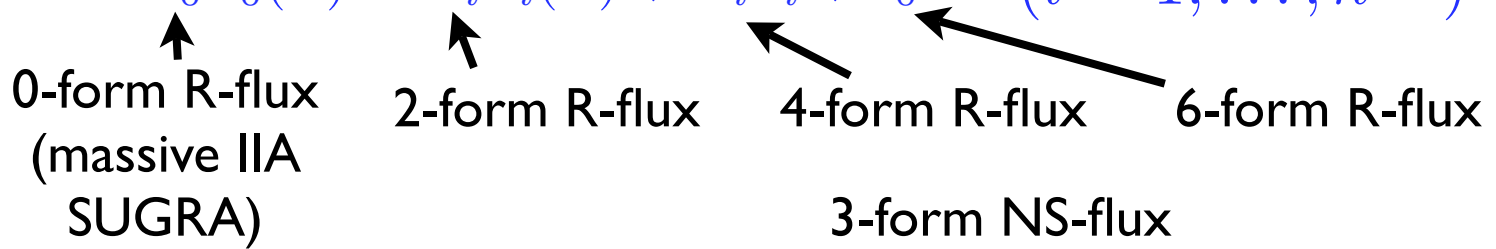


(Curio, Klemm, Lüst, Theisen, hep-th/0012213; Derendinger, Kounnas, Petropoulos, Zwirner, hep-th/0411276; Villadoro, Zwirner, hep-th/0503169; De Wolfe, Giriyavets, Kachru, Taylor, hep-th/0505160; Camara, Font, Ibanez, hep-th/0506066; Villadoro, Zwirner, arXiv:0706.3049)

Consider compactification on CY space Y with Hodge numbers $\tilde{h}^{2,1}$ $\tilde{h}^{1,1}$

Superpotential: $W_{\text{IIA}} = W_H + W_F + W_{\text{geom}}$

$$\begin{aligned}
 W_F(T) &= \int_Y e^{J_c} \wedge F^R \\
 &= \tilde{m}_0 \frac{1}{6} \int_Y (J_c \wedge J_c \wedge J_c) + \frac{1}{2} \int_Y (F_2^R \wedge J_c \wedge J_c) + \int_Y F_4^R \wedge J_c + \int_Y F_6^R \\
 &= i\tilde{m}_0 F_0(T) - \tilde{m}_i F_i(T) + i\tilde{e}_i T_i + \tilde{e}_0. \quad (i = 1, \dots, \tilde{h}^{1,1})
 \end{aligned}$$



$$W_H(S, U) = \int_Y \Omega_c \wedge H_3 = i\tilde{a}_0 S + i\tilde{c}_m U_m, \quad (m = 1, \dots, \tilde{h}^{2,1})$$



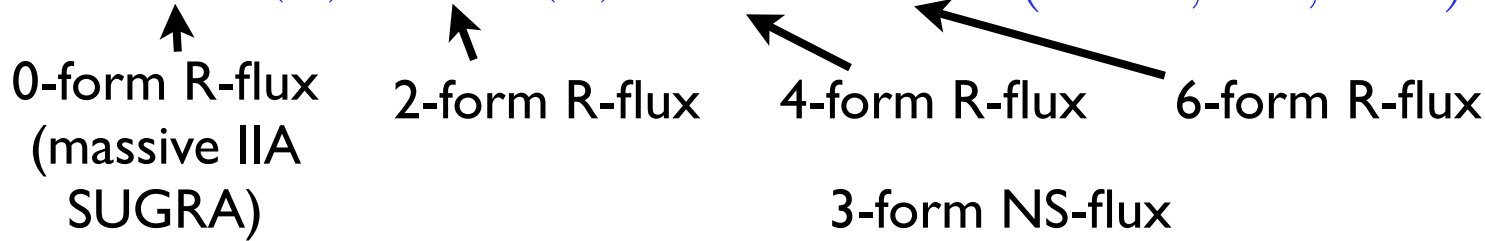
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Metric fluxes (twisted tori)

$$W_{\text{geom}}(S, T, U) = i \int_Y \Omega_c \wedge dJ = -\tilde{a}_i S T_i - \tilde{d}_{im} T_i U_m,$$

The fluxes induce a C7 tadpole:

$$\tilde{N}_{\text{flux}} = \int (C_7 \wedge dF_2 + C_7 \wedge (\tilde{a}_0 H_3 + d\bar{F}_2)) = \sum_{I=0}^{\tilde{h}^{1,1}} \tilde{a}_I \tilde{m}_I .$$

D6-brane charge, need D6-branes, O6-planes.





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($\tilde{h}^{1,1} = \tilde{h}^{2,1} = 3$)



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$$|\gamma_i| T_i = \sqrt{\frac{5|\gamma_1 \gamma_2 \gamma_3|}{3\tilde{m}_0^2}}, \quad S = -\frac{2}{3\tilde{m}_0 \tilde{a}_0} \gamma_i T_i, \quad \tilde{c}_m U_m = -\frac{2}{3\tilde{m}_0} \gamma_i T_i .$$

($\gamma_i = \tilde{m}_0 \tilde{e}_i$)

T-dual type IIB description:



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IIB: No Kähler moduli: (LG description)

(Becker, Becker, Vafa, Walcher, hep-th/0611001)

Then all IIB moduli can be fixed by 3-form flux superpotential:

$$W_{\text{IIB}} = \int_X \Omega \wedge (F_3^{\text{R}} + SH_3^{\text{NS}}) = ie_i U_i + im_0 U_1 U_2 U_3 + ia_0 S$$

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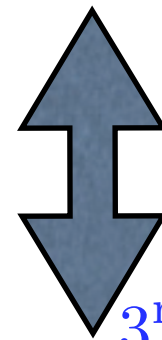
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$$0 = \left(\nabla_{\mu} + \frac{1}{4} H_{\mu} \mathcal{P} + \frac{e^{\Phi}}{16} \sum_n F_n \Gamma_{\mu} \mathcal{P}_n \right) \epsilon ,$$

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- Non-vanishing $C7$ tadpole, same as for fluxes

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Ramond 4-form flux superpotential:

$$W_{\text{IIA}} = \int_Y F_4^R \wedge J = i\tilde{e}_i T_i$$

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The corresponding sources are intersecting D4-branes:

	ξ^0	ξ^1	ξ^2	y^1	y^2	y^3	x^1	x^2	x^3	x^4
D4	⊗	⊗	⊗				⊗	⊗		
D4'	⊗	⊗	⊗						⊗	⊗

$$ds_{10}^2 = \frac{1}{\sqrt{H_1 H_2}} \eta_{\mu\nu} d\xi^\mu d\xi^\nu + \sqrt{H_1 H_2} \delta_{ij} dy^i dy^j + \sqrt{\frac{H_2}{H_1}} \sum_{a=1}^2 (dx^a)^2 + \sqrt{\frac{H_1}{H_2}} \sum_{a=3}^4 (dx^a)^2$$

$$e^{-2\phi} = \sqrt{H_1 H_2}, \quad F_{x^3 x^4 y^i y^j} = -\delta^{kl} \varepsilon_{ijk} \partial_{y^l} H_1, \quad F_{x^1 x^2 y^i y^j} = -\delta^{kl} \varepsilon_{ijk} \partial_{y^l} H_2,$$

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H are not harmonic in overall transverse directions .

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Metric of $\text{AdS}_4 \times S^2 \times T^4 + \sqrt{\frac{D_2}{D_1}} \sum_{a=1}^2 (dx^a)^2 + \sqrt{\frac{D_1}{D_2}} \sum_{a=3}^4 (dx^a)^2$

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String coupling $g_s = e^\phi \rightarrow 0$

Runaway behavior

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$$DW_{\text{IIA}} = 0 \quad \Rightarrow \quad \text{All moduli are stabilized !}$$

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The corresponding sources are intersecting D4, NS5 and D8-branes:

	ξ^0	ξ^1	ξ^2	y	x^1	x^2	x^3	x^4	x^5	x^6
D4	⊗	⊗	⊗		⊗	⊗				
D4'	⊗	⊗	⊗				⊗	⊗		
D4''	⊗	⊗	⊗						⊗	⊗
NS5	⊗	⊗	⊗		⊗		⊗		⊗	
NS5'	⊗	⊗	⊗		⊗			⊗		⊗
NS5''	⊗	⊗	⊗			⊗		⊗	⊗	
NS5'''	⊗	⊗	⊗			⊗	⊗			⊗
D8	⊗	⊗	⊗		⊗	⊗	⊗	⊗	⊗	⊗

Explicit form of the solution:

$$\begin{aligned}
 ds_{10}^2 = & \left\{ H^{D8} \left(\prod_{\alpha=1}^3 H_{\alpha}^{D4} \right) \right\}^{-\frac{1}{2}} \eta_{\mu\nu} d\xi^{\mu} d\xi^{\nu} \\
 & + \left(\prod_{\alpha=1}^4 H_{\alpha}^{NS5} \right) \left\{ H^{D8} \left(\prod_{\alpha=1}^3 H_{\alpha}^{D4} \right) \right\}^{\frac{1}{2}} dy^2 \\
 & + \sqrt{\frac{H_2^{D4} H_3^{D4}}{H_1^{D4} H^{D8}}} \left\{ H_3^{NS5} H_4^{NS5} (dx^1)^2 + H_1^{NS5} H_2^{NS5} (dx^2)^2 \right\} \\
 & + \sqrt{\frac{H_1^{D4} H_3^{D4}}{H_2^{D4} H^{D8}}} \left\{ H_2^{NS5} H_3^{NS5} (dx^3)^2 + H_1^{NS5} H_4^{NS5} (dx^4)^2 \right\} \\
 & + \sqrt{\frac{H_1^{D4} H_2^{D4}}{H_3^{D4} H^{D8}}} \left\{ H_2^{NS5} H_4^{NS5} (dx^5)^2 + H_1^{NS5} H_3^{NS5} (dx^6)^2 \right\};
 \end{aligned}$$

$$e^{2\phi} = \left(\prod_{\alpha=1}^4 H_{\alpha}^{NS5} \right) \left(\prod_{\alpha=1}^3 H_{\alpha}^{D4} \right)^{-\frac{1}{2}} (H^{D8})^{-\frac{5}{2}};$$

$$H_{x^2 x^4 x^6} = -\partial_y H_1^{NS5} (H^{D8})^{-1}; \quad H_{x^2 x^3 x^5} = -\partial_y H_2^{NS5} (H^{D8})^{-1};$$

$$H_{x^1 x^3 x^6} = -\partial_y H_3^{NS5} (H^{D8})^{-1}; \quad H_{x^1 x^4 x^5} = -\partial_y H_4^{NS5} (H^{D8})^{-1};$$

$$F_{x^3 x^4 x^5 x^6} = \partial_y H_1^{D4}; \quad F_{x^1 x^2 x^5 x^6} = \partial_y H_2^{D4};$$

$$F_{x^1 x^2 x^3 x^4} = \partial_y H_3^{D4}; \quad F = -\partial_y H^{D8} \left(\prod_{\alpha=1}^4 H_{\alpha}^{NS5} \right)^{-1}.$$

Properties of this solution:



- Smearred (thick) branes:

$$H_{\alpha}^{\text{D4}} = \begin{cases} c_{\alpha}^{\text{D4}} y \left\{ 1 - \frac{1}{2} \left(\frac{y}{y_0} \right) \right\}, & y < y_0 \\ \frac{1}{2} c_{\alpha}^{\text{D4}} y_0, & y \geq y_0 \end{cases}$$

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- Non-vanishing tadpole: need (smearred) D6-branes and O6-plane.

T-dual type II B description:



T-duality in x_1 direction:

$$W_{\text{IIB}} = i(\tilde{e}_1 U_1 + \tilde{c}_2 U_2 + \tilde{c}_3 U_3 + \tilde{c}_1 T_1 + \tilde{e}_2 T_2 + \tilde{e}_3 T_3 + \tilde{m}_0 U_1 T_2 T_3 + \tilde{a}_0 S)$$

This includes geometrical fluxes.

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Corresponding branes:

	ξ^0	ξ^1	ξ^2	y	x^1	x^2	x^3	x^4	x^5	x^6
D3	⊗	⊗	⊗			⊗				
D5	⊗	⊗	⊗		⊗		⊗	⊗		
D5'	⊗	⊗	⊗		⊗				⊗	⊗
D7	⊗	⊗	⊗			⊗	⊗	⊗	⊗	⊗
NS5	⊗	⊗	⊗		⊗		⊗		⊗	
NS5'	⊗	⊗	⊗		⊗			⊗		⊗
KK	⊗	⊗	⊗		•	⊗		⊗	⊗	
KK'	⊗	⊗	⊗		•	⊗	⊗			⊗

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D5'	⊗	⊗	⊗		⊗				⊗	⊗
D7	⊗	⊗	⊗			⊗	⊗	⊗	⊗	⊗
NS5	⊗	⊗	⊗		⊗		⊗		⊗	
NS5'	⊗	⊗	⊗		⊗			⊗		⊗
KK	⊗	⊗	⊗		•	⊗		⊗	⊗	
KK'	⊗	⊗	⊗		•	⊗	⊗			⊗

Near horizon: $AdS_4 \times N$, (Nilmanifold)

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Open problems:

- Entropy count of brane configurations.
- Uplift to dS and its brane description.
- Characterization of brane configurations via generalized geometry.

Conclusions



How does the string landscape really look like?



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