

Using GPUs for LSS analysis

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Speeding up (your analysis) with a graphics card

Summary / Short version of the talk

- **GPUs** are **inexpensive** and **powerful** tools for **2-point correlation function** computation.
- The **code** to profit from this is **available**:

<http://wwwae.ciemat.es/cosmo/gp2pcf>

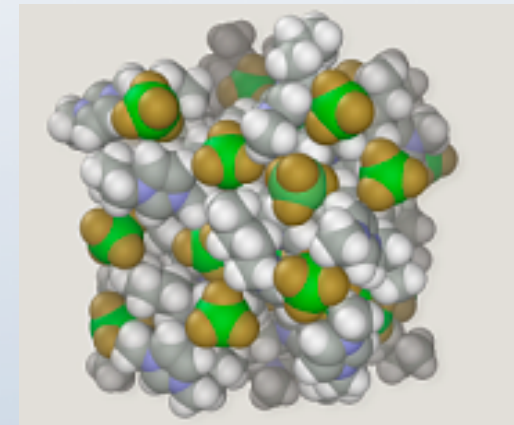
GPUs have been used for a long time

Graphics Processing Unit:

Originally: 2D-3D rendering (1980's - today)

Floating point operations, high parallelization

Now also physics simulations, **scientific computing** (GPGPU)

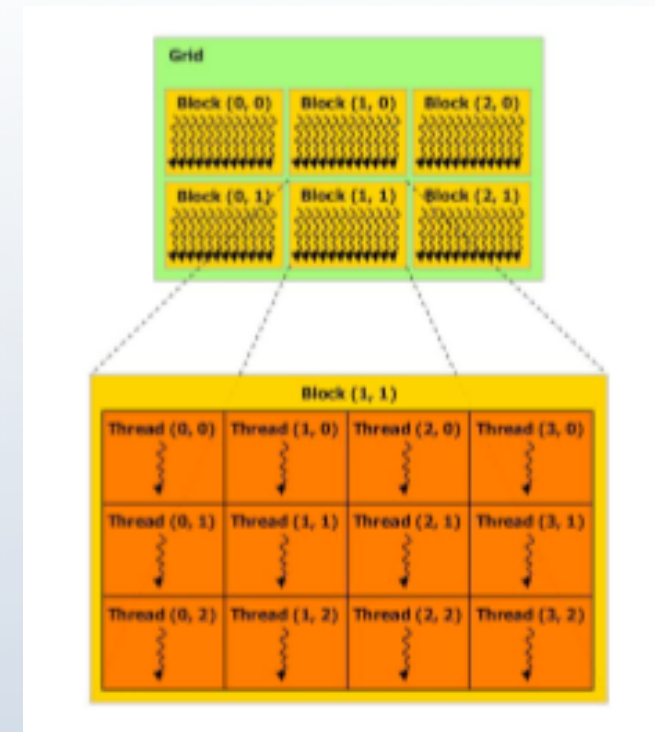


GPUs are designed specifically for parallel processing

Nowadays usually on separate card.

Most of the transistors devoted to operations.

Several blocks are sent, shared between the GPU's processors; each block: hundreds of threads doing parallel operations.

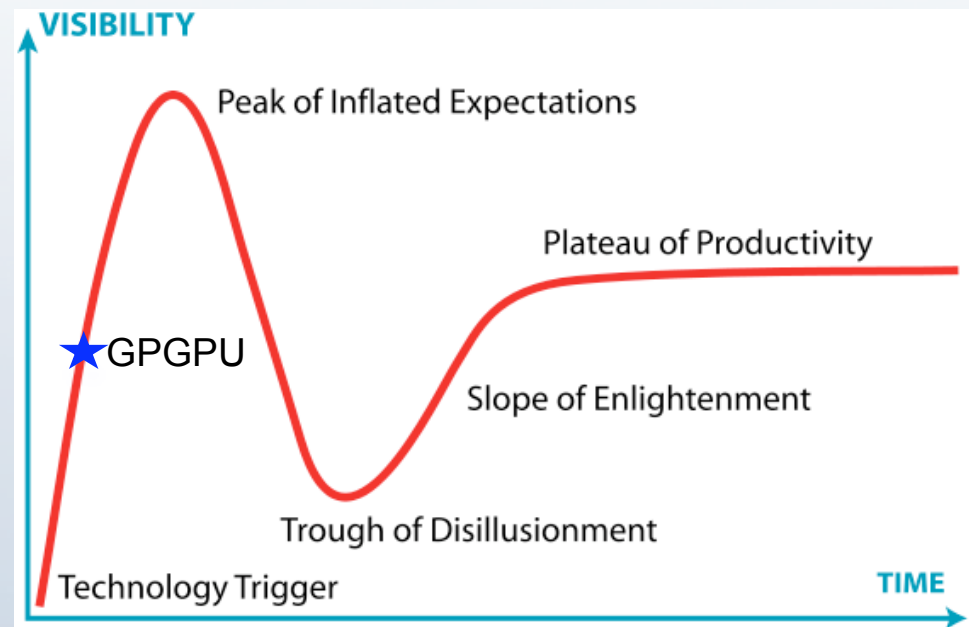


Credit: T. Jensen, H.Rasmussen, F. Rosé

Only very recently they are starting to be used beyond graphics processing

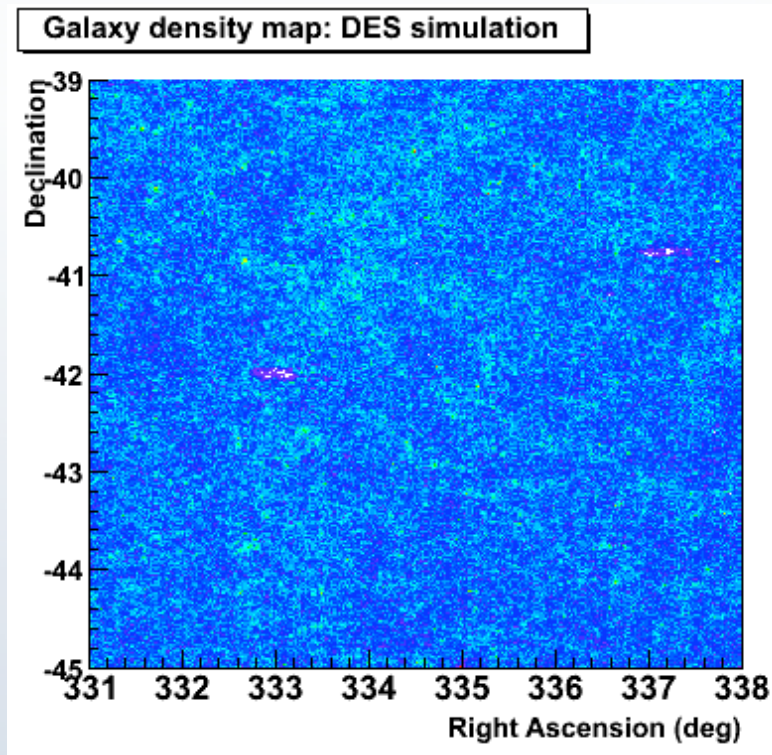
Multiple applications beyond graphics (GPGPU)

Only very specific applications (massive, parallel)

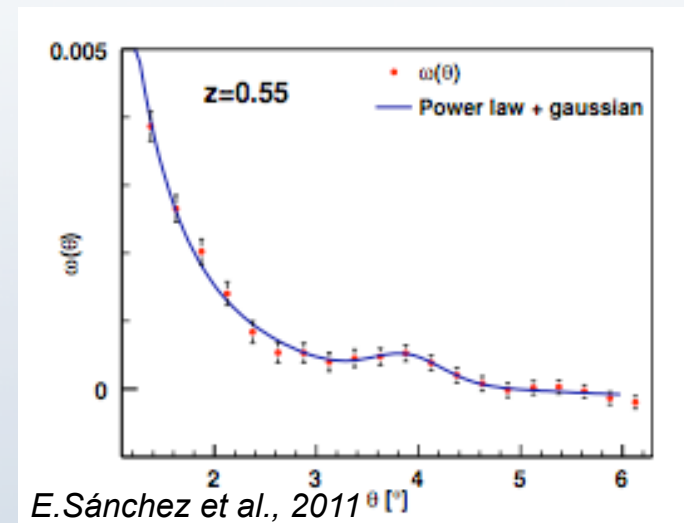


Gartner hype plot, adapted

Cosmology can benefit from this device: e.g. the angular 2pc function computation



Measure **deviation** of distribution from **random**.
Shape, normalization, features -> **cosmological interpretation**.



E.Sánchez et al., 2011

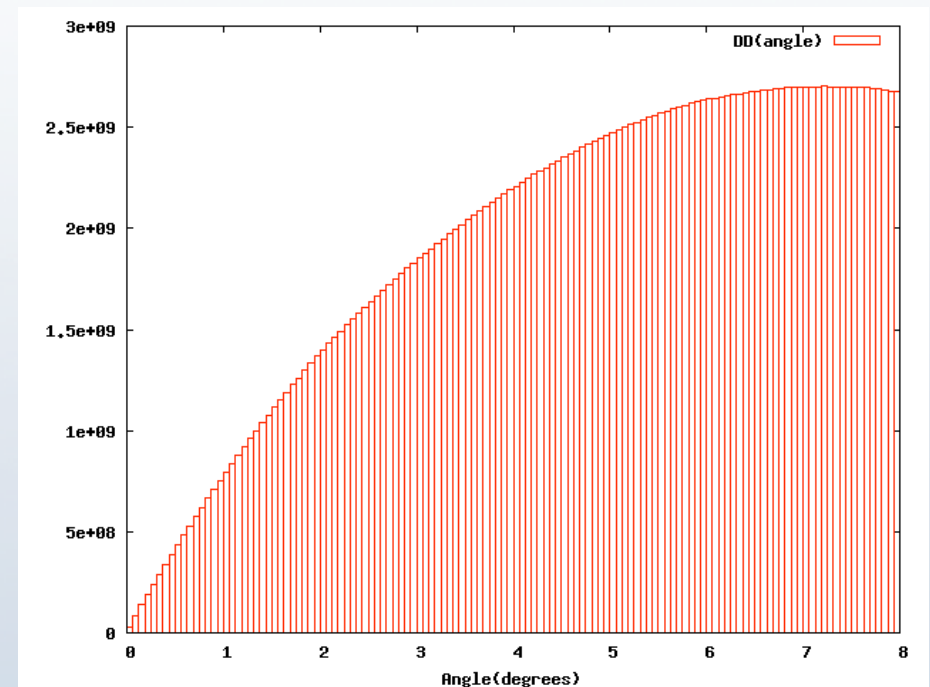
$$\omega(\theta) = 1 + \left(\frac{N_{\text{random}}}{N_{\text{real}}}\right)^2 \cdot \frac{DD(\theta)}{RR(\theta)} - 2 \cdot \left(\frac{N_{\text{random}}}{N_{\text{real}}}\right) \cdot \frac{DR(\theta)}{RR(\theta)}$$

Landy & Szalay, 1993

We have implemented the full angular 2pc function algorithm in CUDA

CUDA is a C extension to interact with GPU using C-like syntax.

- Kernel: C-like function (*Compute angle between two galaxies*)
- N threads: execute kernel N times in parallel.
- Blocks: threads grouped into blocks (*DD, DR, RR histos built in each block*).



Designing the kernel needs to specially consider the features of the hardware

```
void binning(float *xd, float *yd, float *zd,
            int *ZZ, int numberoflines){

    float angle;
    int dibin[angleposition];

    for (int i=0; i< numberoflines; i++){
        for (int j=0; j< numberoflines; j++){
            angle = xd[i]*xd[j] + yd[i]*yd[j] + zd[i]*zd[j];
            if(angle>1.) angle=1.;
            angle = acosf(angle) * 180./pi;
            position = int((angle)*binsperdegree);
            if(dibin[position]<binsperdegree*totaldegrees);
                dibin[position] = dibin[position]+1;
        }
    }
}
```

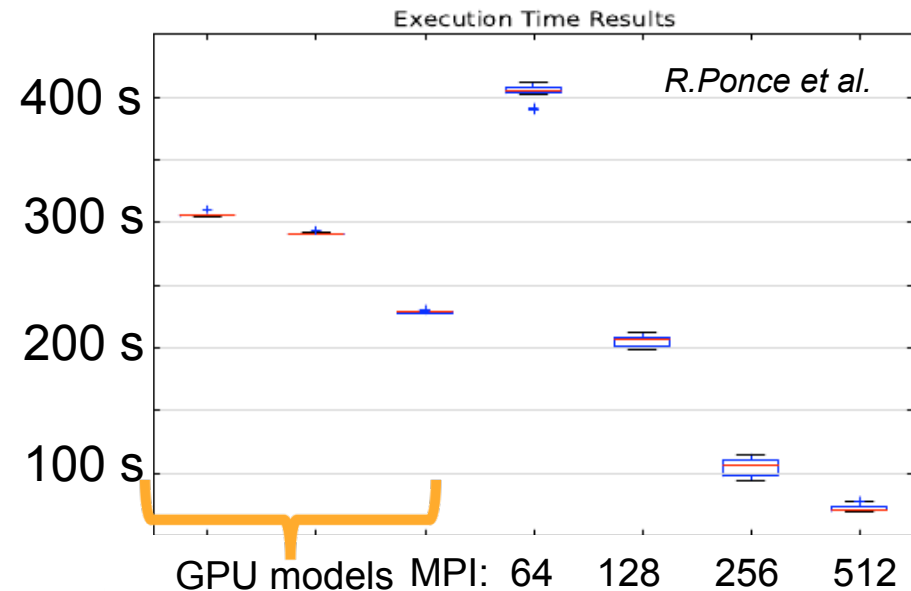
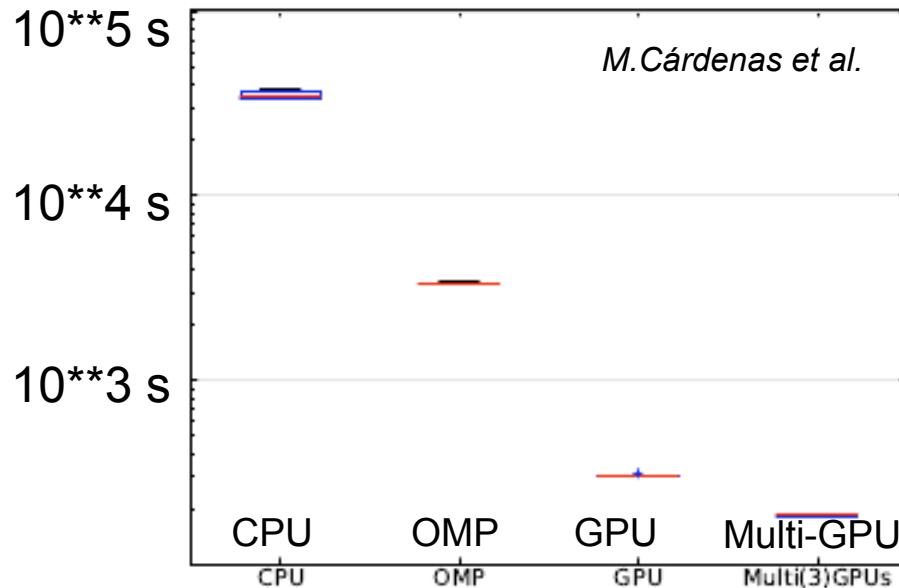
CPU Code

```
__global__ void binning(float *xd, float *yd, float *zd,
                        int *ZZ, int numberoflines){
    __shared__ float angle[threadsperblock];
    __shared__ int dibin[threadsperblock];
    __shared__ int temp[threadsperblock];
    int cacheIndex = threadIdx.x;
    temp[cacheIndex]=0;
    for (int i=0; i< numberoflines; i++){
        int dim_idx = blockIdx.x * blockDim.x + threadIdx.x;
        while (dim_idx < numberoflines){
            angle[cacheIndex] = xd[i]*xd[dim_idx] + yd[i]*yd[dim_idx] +
                                zd[i]*zd[dim_idx];
            __syncthreads();
            if(angle[cacheIndex]>1.) angle[cacheIndex]=1.;
            .
            .
            .
            dim_idx += blockDim.x * gridDim.x; } }
    atomicAdd( &ZZ[threadIdx.x] , temp[threadIdx.x]); }
```

GPU Code

A for-loop is substituted by parallel computation

Results: single GPU is equivalent to ~100 processors



CPU: Intel Xeon 4 cores @ 2.27 GHz

OMP and MPI: Supercomputer with 240 nodes x 8 cores @ 2.93 GHz

GPU: 3 different models

Multi-GPU: 3 x GTX295 model dividing the work of DD, DR, RR

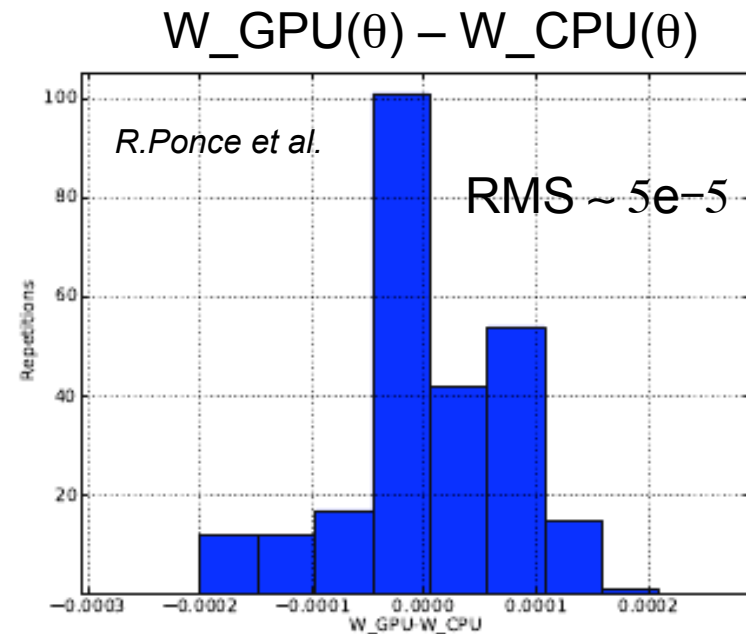
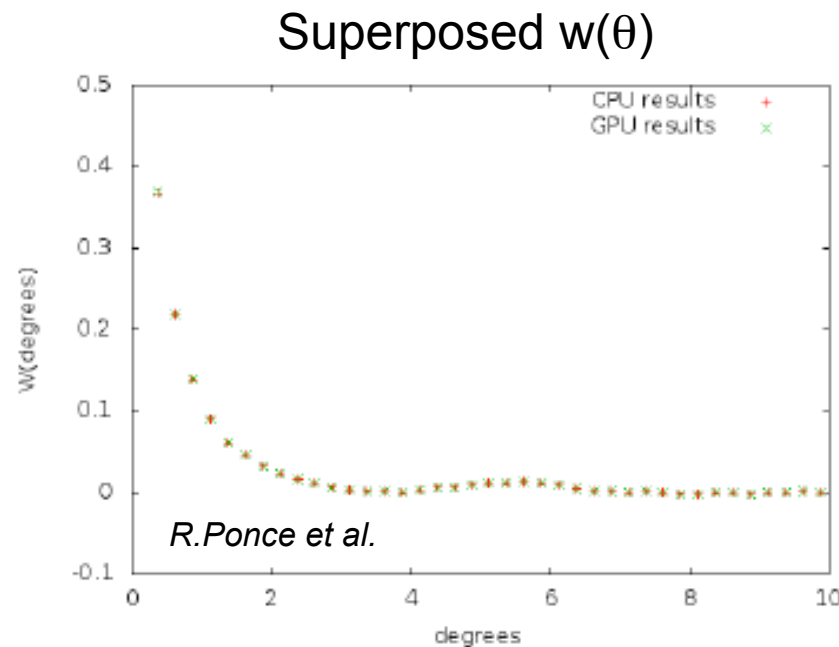
Results: a million galaxies in 20 minutes

Input file lines	CPU (s)	GTX295 (s)	C1060 (s)	C2050 (s)
$0.43 \cdot 10^6$	$3.60 \cdot 10^4$	$3.01 \cdot 10^2$	$2.91 \cdot 10^2$	$2.19 \cdot 10^2$
$0.86 \cdot 10^6$	$1.44 \cdot 10^5$	$1.20 \cdot 10^3$	$1.16 \cdot 10^3$	$8.76 \cdot 10^2$
$1.00 \cdot 10^6$	$1.98 \cdot 10^5$	$1.61 \cdot 10^3$	$1.56 \cdot 10^3$	$1.17 \cdot 10^3$
$1.29 \cdot 10^6$	$3.24 \cdot 10^5$	$2.68 \cdot 10^3$	$2.59 \cdot 10^3$	$1.97 \cdot 10^3$
$1.72 \cdot 10^6$	$5.76 \cdot 10^5$	————	$4.64 \cdot 10^3$	$3.51 \cdot 10^3$
$3.45 \cdot 10^6$	$2.32 \cdot 10^6$	————	$1.88 \cdot 10^4$	$1.41 \cdot 10^4$
$6.89 \cdot 10^6$	$9.22 \cdot 10^6$	————	$7.45 \cdot 10^4$	$5.61 \cdot 10^4$

M.Cárdenas et al.

Intel Xeon 4-cores against 3 GPU models

Results: residuals negligible wrt statistical errors



Residuals are smaller using latest GPU cards (single $\rightarrow$ double precision)

There are some constraints when developing for GPUs

- Steep learning curve (not so if used to parallel processing)
- Applications limited by memory available at GPU
- Binning constraints

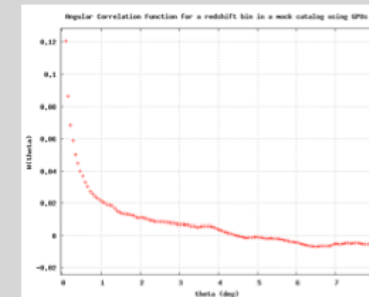
Presenting: GP2pcf

We provide source code in CUDA, documentation and support.

Tarball with code and examples available at:

<http://wwwae.ciemat.es/cosmo/gp2pcf>

FAST 2 POINT CORRELATION FUNCTIONS USING GPUS



Angular correlation function computed with GPUs on a mock catalog.

GP2PCF: a code for brute-force computation of 2 point correlation functions

Authors

- Miguel Cárdenas-Montes
- Rafael Ponce

Abstract

The two-point correlation function is a simple statistic that quantifies the clustering of a given distribution of objects. In studies of the large scale structure of the Universe, it is an important tool containing information about the matter clustering and the evolution of the Universe at different cosmological epochs. A classical application of this statistic is the galaxy-galaxy correlation function to find constraints on the parameter Ω_m or the

Notes on the implementation available

A preprint describing the code and tests has been submitted to Computer Physics Communications. **Available on request!**

Also check out XXI ADASS proceedings (directly available for download from GP2PCF website).

Performance Study for Parallel Implementations of Cosmological Data Analysis

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Abstract

The rapid increase of the amount and quality of observations on the large scale structure of the Universe has opened a new epoch in cosmology. Simulations and data analysis are performed more and more frequently in diverse computational platforms: GPU, clusters or supercomputers. In this work, the performance of these platforms is tested through the calculation of the Two-Point Angular Correlation Function of galaxies, that allows to select the most suitable for the coming production phase. The Two-Point Angular Correlation Function is one of the tools commonly used to test the Λ CDM model through measurements of large-scale structures, which is a relevant topic in cosmology. Furthermore, taking into account the high volume of observational data expected in the coming years, any improvement in the execution time of this particular analysis is welcome. In this work, some computational platforms are tested; namely GPU and mid-size supercomputer. Moreover,

Preprint submitted to Computer Physics Communications

November 17, 2011

Future developments

- k-tree instead of brute-force approach
- galaxy-shear correlation
- N-point correlation function
- **3D correlation function**
- **multi-GPU**
- **pyCUDA version**

Other uses in cosmology

- Matching catalogs
- K-nearest neighbors (e.g. photo-z) (V. García)
- N-body simulations (C. Frigaard)
- ...

Summary



GPUs are **inexpensive** and **powerful** tools for **2-point correlation function** computation. The cost is lower in equivalent chips, Watts/Gflop, personnel, fight for resources.



The **code** to profit from this is **available**:

<http://wwwae.ciemat.es/cosmo/gp2pcf>

Summary



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“Supercomputing for the masses” Dr. Dobbs

GFLOPS

