Hawking Radiation in Acoustic Black-Holes on an Ion Ring

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Towards the Observation of Hawking Radiation in Condensed Matter Systems. IFIC, Valencia, February 1-7, 2009

Outline:

1. A discrete analogue model of a sonic hole toy model on a ion ring geometry.

-- motivation, trapped ions.

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- --numerical classical simulation of a BH.
- 2. Numerical evidence of Hawking radiation:
 - -- backward evolution and Bogolibouv particle creation.
 - --mode conversion (Unruh's sonic-hole)
 - --Bloch oscillations (Corley and Jacobson, "falling lattice")
 - --Dynamical creation and point-to-point correlations.

Sonic dumb-hole

Unruh 81, 95



Ion lattices



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Oxford, England: 40Ca+





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Innsbruck, Austria: ⁴⁰Ca⁺





Aarhus, Denmark: ⁴⁰Ca⁺ (red) and ²⁴Mg⁺ (blue)





Trapped ions



Size of the wave packet << wavelength of visible light.

Entanglement entropy in Field Theory



Could it be that Entanglement is the quantum source of BH entropy?

Bombelli, Koul, Lee, Sorkin 86

Entanglement entropy

2-D Harmonic lattice



 $\text{Entanglement} \propto \text{ Area}$

$$S_{AB} = -tr(\rho_A \log \rho_A)$$

Simulating detection of vacuum entanglement



Entanglement Entropy

 $H=H_0+H_{int}$

$$H_0 = \omega_z (\sigma_z^A + \sigma_z^B) + \sum v_n a_n^\dagger a_n$$

$$H_{int} = \Omega(t) (e^{-i\phi}\sigma_{+}^{(k)} + e^{i\phi}\sigma_{-}^{(k)}) x_{k}$$

 $1/\omega_z << T << 1/v_0$

Discrete BH analogue



Harmonic oscillations around the equilibrium motion are phonons with velocities c (θ) \propto (v (θ))-1/2.

When v increases the sound velocity decreases and a Black and White horizons can form.

Observation of Ordered Structures of Laser-Cooled Ions in a Quadrupole Storage Ring

I. Waki,^(a) S. Kassner, G. Birkl, and H. Walther Nature (1992), PRL (1992)



lons on a ring

N ions of mass m in a ring with radius L:

$$\mathcal{H} = -\sum_{i=1}^{N} \frac{4\pi^2 \hbar^2}{2mL^2} \frac{\partial^2}{\partial \theta_i^2} + \sum_{i=1}^{N} V^e\left(\theta_i\right) + V^c\left(\theta_1, \dots, \theta_N\right)$$

We treat small perturbations around the equilibrium motion

$$\theta_{i}(t) = \theta_{i}^{0}(t) + \delta\theta_{i}(t)$$

and expand the Hamiltonian to second order in $\delta \theta_i$

$$\mathcal{H} = -\sum_{i=1}^{N} \frac{4\pi^2 \hbar^2}{2mL^2} \frac{\partial^2}{\partial \delta \theta_i^2} + \frac{1}{2} \sum_{i \neq j} f_{ij}(t) \delta \theta_i \delta \theta_j \tag{1}$$

Large N effective field limit

For a slowly varying $v(\theta)$ the system Lagrangian for the scalar field $\Phi(\theta_i^0(t), t) = \delta\theta_i(t)$ becomes

$$\mathcal{L} = \frac{mL^2}{(2\pi)^2} \int d\theta \frac{n\left(\theta\right)}{2} \left[\left(\partial_t \Phi + v\left(\theta\right) \partial_\theta \Phi\right)^2 - \left(iD\left(\theta, -i\partial_\theta\right) \Phi\right)^2 \right]$$

with $D(\theta, k) = c(\theta) k + \mathcal{O}(k^3)$, the density $n(\theta) = 1/(v(\theta)T)$. $c(\theta) = \sqrt{2(2\pi)^3 n(\theta) e^2/(4\pi\epsilon_0 \cdot mL^3)}$.

We have a position dependent dispersive fluid.

Creating a discrete BH

We want the classical dumb-hole equilibrium motion:

$$\theta_i(t) = g_v(\frac{i+t/T}{N})$$

where g maps the normalized indices i/N \in [0,1] monotonically increasing onto the angles $\theta \in$ [0,2 π] and is periodically continued.

The velocity profile is then $v(\theta) = g'(g^{-1}(\theta))/T$

Dynamics

$$\frac{\partial}{\partial t} \langle \hat{\xi}_i \rangle_t = \sum_j G_{ij}(t) \langle \hat{\xi}_j \rangle_t$$

$$\frac{\partial}{\partial t}\Gamma(t) = G(t)\cdot\Gamma(t) + \Gamma(t)\cdot G(t)^{T}.$$

Note that since the initial state is Gaussian (either thermal or vacuum) The state will remain Gaussian for later times. Hence higher order Correlations can be computed via Wick's decomposition theorem.

Group velocity



For only nearest neighbor interactions $c = 2e\sqrt{N}$.

or
$$c\left(\theta\right) = \sqrt{2(2\pi)^3 n\left(\theta\right) e^2 / (4\pi\epsilon_0 \cdot mL^3)}$$

Testing wave packet trajectories

From Unruh's model we expect:

Dispersion relation: $\omega = F(k)$ (fluid rest frame)



-reflection at the horizon

-Hawking's radiation thermal properties remain mostly unaffected -but in our case k has a maximal value.

Back in time

$$\delta \theta_k^n(0) = k \cdot e^{-(\frac{k-2\pi n^2}{40\pi})}, \quad n = 1,...,20,$$

Outgoing wave (negative frequency)

Values of $\delta \theta_i(t)$ during propagation backwarts in time. Here we have v = 0.83, k = 10, N = 1000, N2 = 1200,



Scattering to Low k modes

We find scattering to a low wavenumber negative frequency mode, on the second branch, which is mostly right moving. This scattering is consistent with the approximate Killing frequency conservation.

It was argued that conformal invariance in 1D prevents scattering. We find that both for the full and truncated models there is a small nonvanishing scattering.

In the commoving frame





Wave remains a right mover

Scattering to a left mover

Unruh's process

Corley & Jacobson's Bloch oscillation



Late time outgoing and early time ingoing distributions $|\Delta x_i(t)|$ in the lab frame.

Positive and negative frequencies



Positive and negative modes

If the phononic excitations are localized in the flat subsonic region, the excitations $\delta\theta_i(t) = \langle \delta\theta_i \rangle_t$ and $\delta\dot{\theta}_i(t) = \langle -i\hbar\partial_{\delta\theta_i} \rangle_t$ can be expressed as modes $\delta\theta_k(t)$ and $\delta\dot{\theta}_k(t)$ with wavenumber k. The positive and negative frequency part of these excitations are defined by

$$\begin{split} \delta\theta_k^{\pm}(t) &= \frac{1}{2} \left(\delta\theta_k(t) \pm i\delta\dot{\theta}_k(t)/\omega_k \right), \\ \delta\dot{\theta}_k^{\pm}(t) &= \frac{1}{2} \left(\delta\dot{\theta}_k(t) \mp i\omega_k\delta\theta_k(t) \right). \end{split}$$

Norm distribution: $\mathcal{N}_{k}^{\pm} = \delta \theta_{k}^{\pm} \delta \theta_{k}^{\pm} - \delta \theta_{k}^{\pm*} \delta \theta_{k}^{\pm}$

First mechanism: mode conversion



Late time (final) negative frequency pulse is depicted in Green. Initial negative frequency pulse in light blue (dashdotted), and positive frequency pulse in red (dashed).

In the commoving frame



Wave remains a right mover

Unruh's process

Second mechanism: Block-like oscillations (Jacobson 98)

Soltions to $\omega_0 = vk \pm D(k)$ lie outside the Brillouin zone:



In the commoving frame



Scattering from right to left mover

Testing Hawking's hypothesis

We compare between the occupation number of the outcoming Wave packet under the thermal Hypothesis And the the Bogoliubov coefficient:

$$\beta_T = \sum_k \frac{\mathcal{N}_k^-(final)}{e^{\frac{\hbar\omega_k}{k_BT}} - 1} = \sum_k \mathcal{N}_k^+(initial)$$

For different late timewaves: $\delta \theta_k^n(0) = k \cdot e^{-(\frac{k-2\pi n^2}{40\pi})}, \quad n = 1,...,20,$

Which taken to be localized at flat space.

Hawking temperature

The Black Hole horizon is located at $c(\theta_H)=v(\theta_H)$. The temperature T_H of Hawking radiation is given by the expression (c depends on θ)

$$\frac{k_B T_H}{\hbar} = -\frac{1}{4\pi v} \frac{d}{d\theta} (c^2 - v^2) |_H = \frac{1}{2\pi} \frac{d}{d\theta} (v - c) |_H.$$

In the case of a Coulomb chain with nearest neighbor interactions only, where the sound velocity of an homogeneous system $c=\$, it can be evaluated in the local density approximation as

$$\frac{k_B T_H}{\hbar} = \frac{3}{4\pi} \frac{g''(g^{-1}(\theta))}{g'(g^{-1}(\theta))}\Big|_{\theta=\theta_H}$$

Numerical results

1. If only nearest neighbor Coulomb interactions are considered, the relative difference between these quantities is lower than ϵ <0.01 in our calculations with up to N=1000 ions.

2. For the long range Coulomb interactions it is of the order of ϵ <0.14. Here the definition of a global Hawking temperature is difficult because of the non-linear dispersion relation at low wavenumbers. Possibly also due to the "non-conformal" scattering.

3. We have checked, that anharmonic effects do not significantly alter these results for oscillation amplitudes comparable to $\langle \theta^2 \rangle$ at the Hawking temperature.

Therefore qualitatively Hawking radiation (mode mixing) seems to persists even in a fundamentally discrete system with long range interactions and a logarithmically diverging group velocity at low wavenumbers

Experimental Parameters

The ion velocity must lie in the same order of magnitude as the phonon velocity in the proposed experimental setup.

For N=1000 singly charged ⁷Li ions with an average spacing of L/N=10 μ m the rotation frequency of the ions must be ω_{ion} =6.3kHz. The Hawking temperature in this system is $k_B T_H / \hbar = 9.8 \omega_{ion} = \underline{62 \text{ kHz}}$.

These parameters and such a temperature can be realized experimentally.

Creation of a black hole

Experimental sequence to measure evidence for Hawking radiation on ion rings:

Begin with a thermal state of the excitations around homogeneously spaced subsonic rotating ions.

Create in a short time a supersonic region to avoid white hole and finite size effects. We change the velocity profile parameter in an exponentially smooth way.

In our case we used
$$0.01 < rac{T_{creation}}{T} < 0.1$$

So we have a time of order up to 0.5 T to observe the radiation.

in Stability



Deviation of the ions position relative to equilibrium as function Of time. BH Creation time is here ~ 0.01 .

Correlations

We analyze the correlations obtained from the covariance matrix \varGamma

Local creation operators:

$$a_{i}^{\dagger} = \frac{1}{\sqrt{2\hbar}} (f_{ii}^{1/4} \delta \theta_{i} - i f_{ii}^{-1/4} \delta p_{i})$$
$$C_{ij}(t) = \frac{\langle (\hat{n}_{i} - \langle \hat{n}_{i} \rangle)(\hat{n}_{j} - \langle \hat{n}_{i} \rangle) \rangle}{\langle \hat{n}_{i} \rangle \langle \hat{n}_{j} \rangle}$$

Where we use Wick's theorem:

$$\langle a_i^{\dagger} a_j^{\dagger} a_i a_j \rangle = \langle a_i^{\dagger} a_i \rangle \langle a_j^{\dagger} a_j \rangle + \langle a_i^{\dagger} a_j^{\dagger} \rangle \langle a_i a_j \rangle + \langle a_i^{\dagger} a_j \rangle \langle a_j^{\dagger} a_i \rangle$$

Note that the interpretation is quite different compared to the BEC case!

Density-density correlations t=0,0.1,0.2,0.3,0.4T R=2.25





On the left Non-normalized correlations. t=0.4T Possible breakdown of the harmonic approximation.

Testing kinematics by modifying the charge e²



The Black hole is not created.

Different in/out relative velocities

Summary

(a) We suggest another avenue towards realization of an analogue BH with cold ions.

(b) This fully discrete Physical analogue model agrees with previously suggested theoretical models.

(c) It appears to be accessible in today's experiments with ions.

(d) Various other tests can be easily done in this model.