

# ***a Bose-Einstein condensate with tunable interaction***

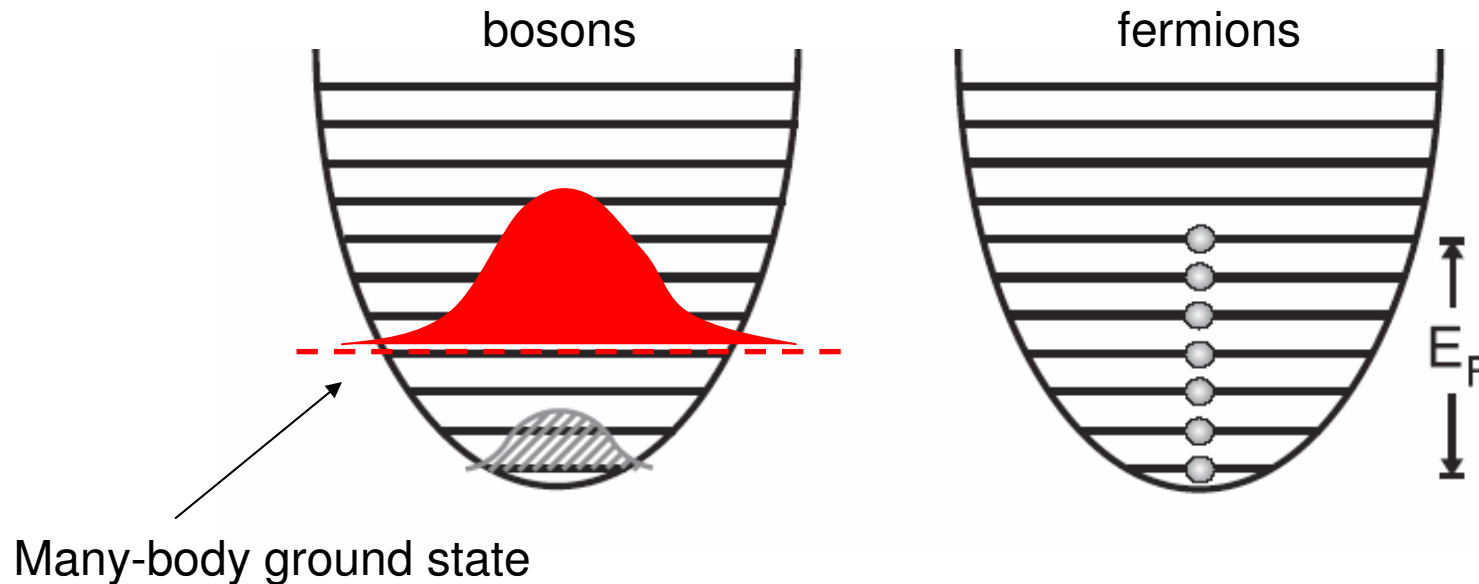
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EHR Workshop, Valencia, February 2009



# Weakly interacting BEC



The interaction energy cannot be neglected:  $E = K + U + E_{\text{int}}$

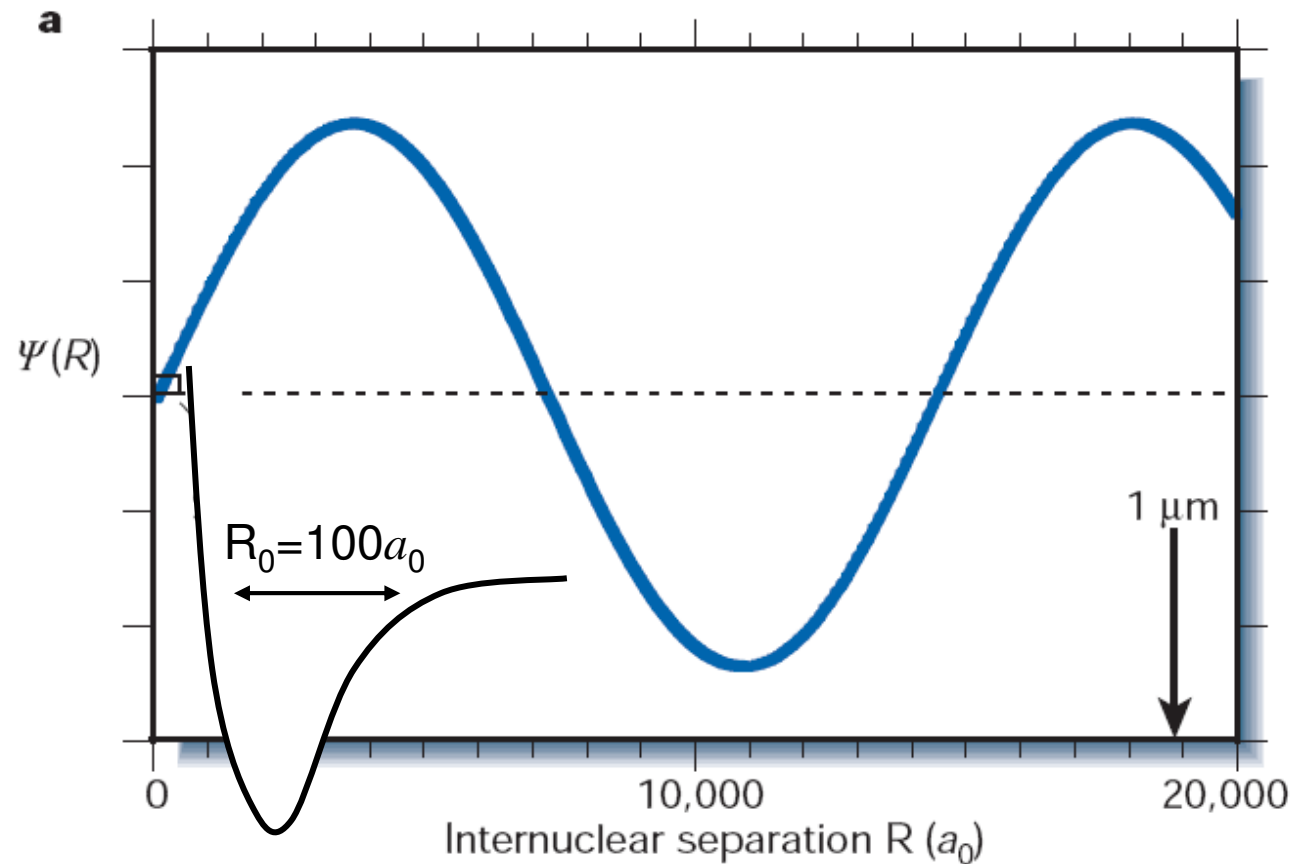
$$E_{\text{int}} = gn \qquad g = \frac{4\pi\hbar^2}{m} a$$

Interaction makes the BEC superfluid and interesting

Interaction can be tuned, turned from attractive to repulsive and in some cases almost cancelled

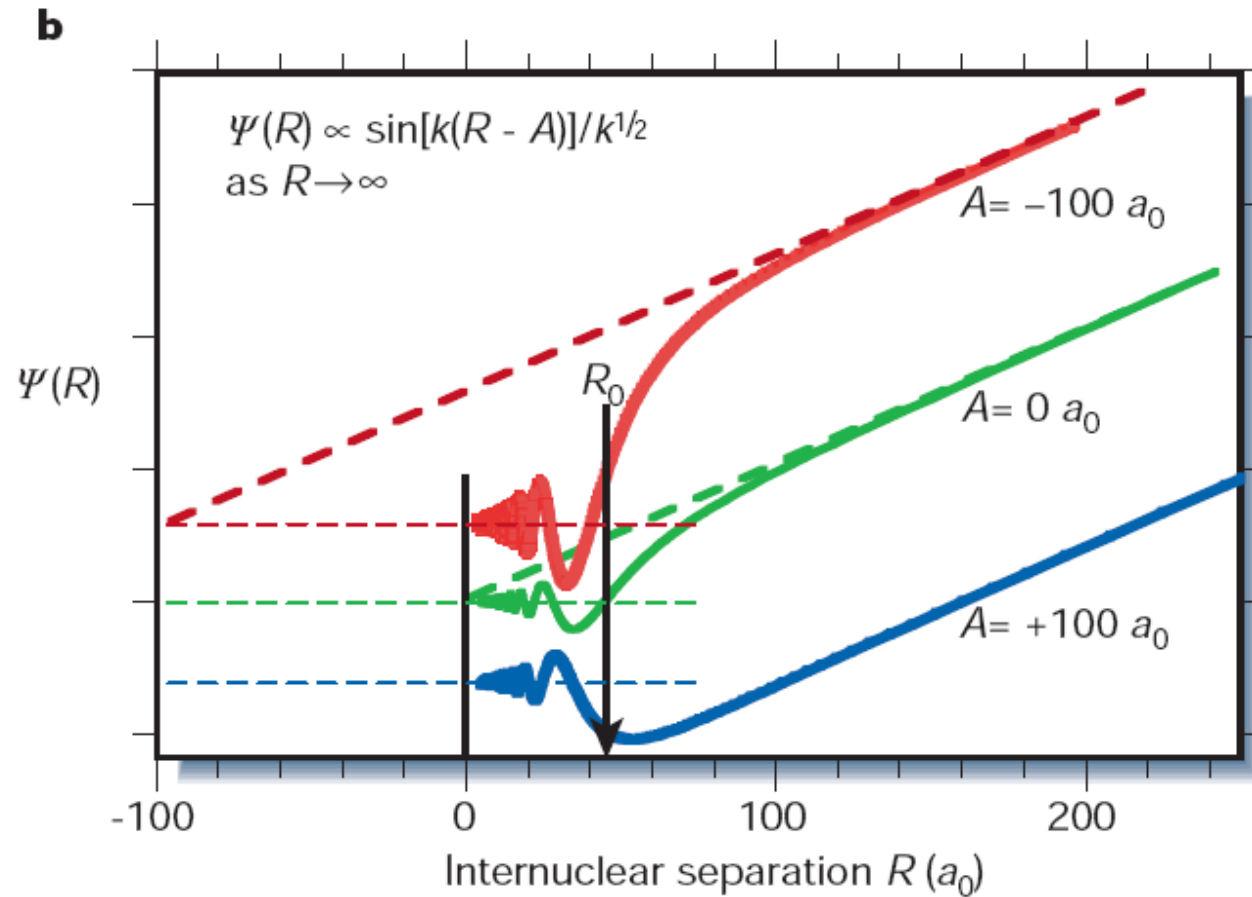
# Scattering length

Ultracold atoms interact via a two-body isotropic contact potential



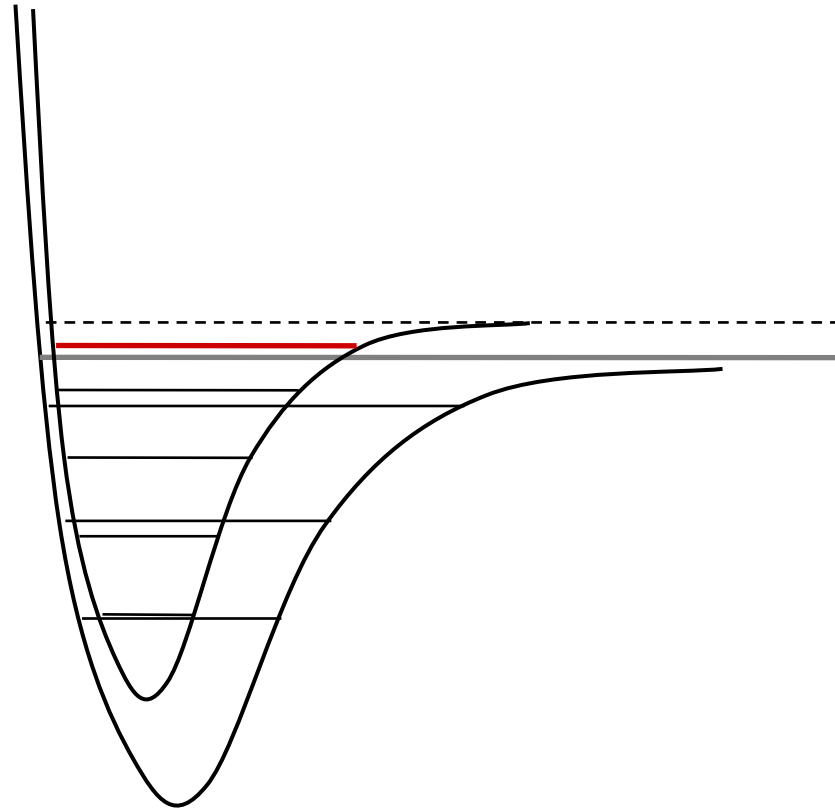
# Scattering length

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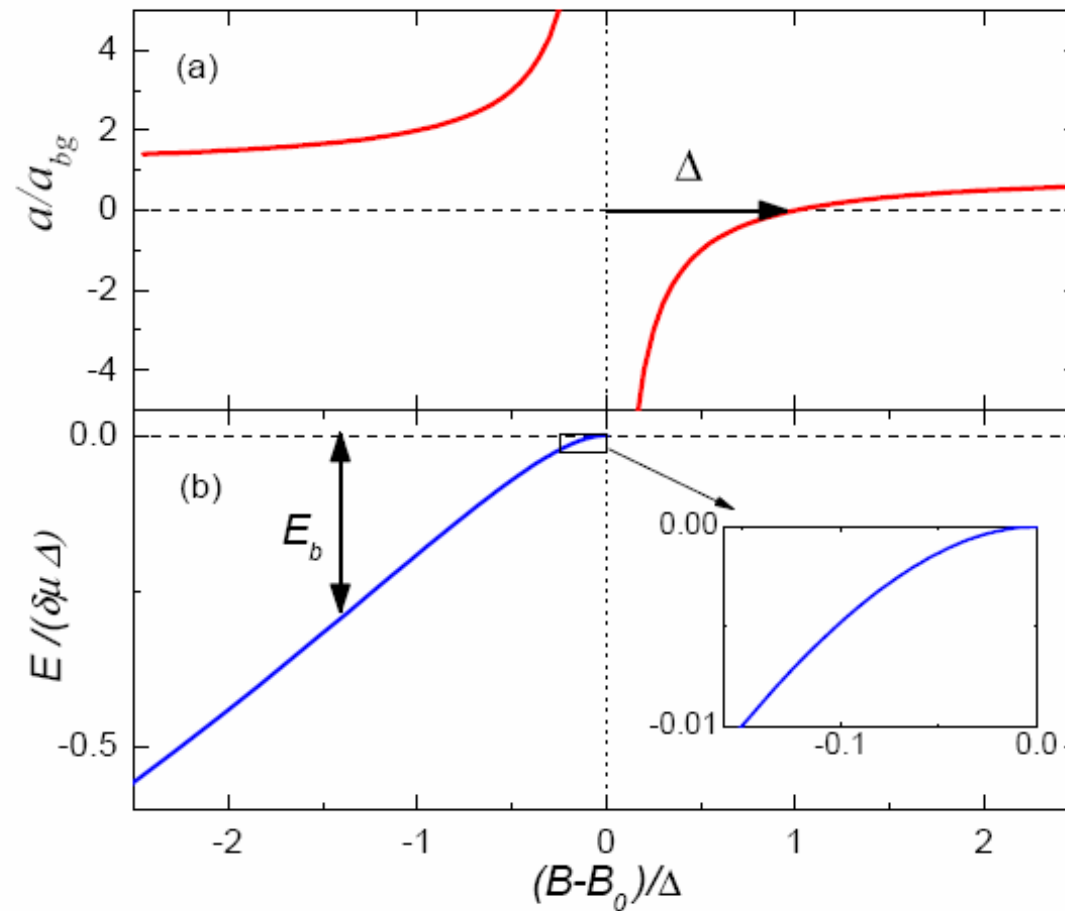
# Scattering length

The scattering length depends on details of the molecular potential (the position of the last bound state)



If the state has an internal structure and a magnetic moment,  $a$  can be varied

# Feshbach resonances

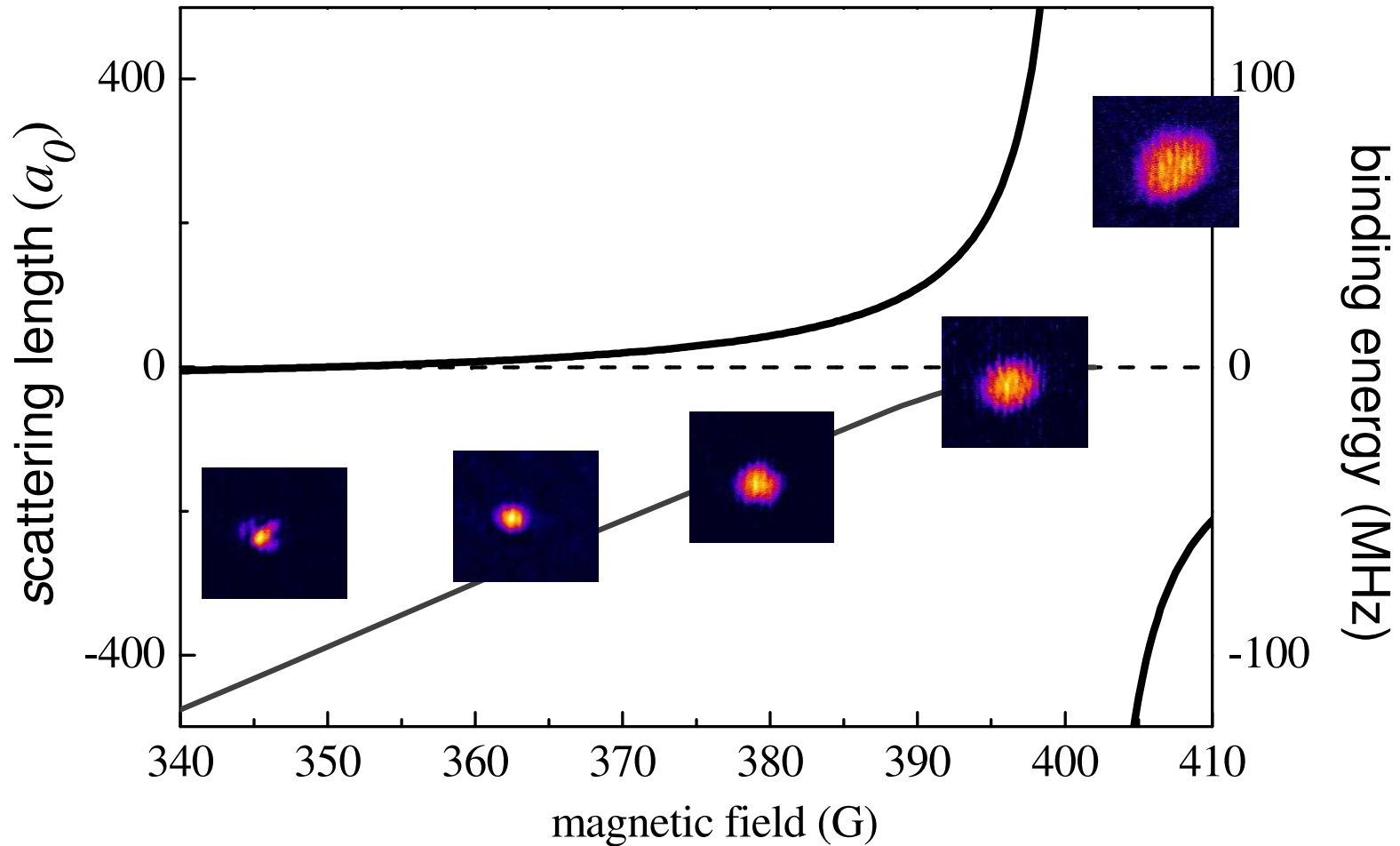


Original idea: H. Feshbach, Ann. Phys. (NY) 5, 357 (1958), U. Fano, Phys. Rev. A 124, 1866 (1961).

Ultracold gases: Moerdijk et al, Phys. Rev. A 51, 4852 (1995); Inouye et al, Nature 392, 151 (1998).

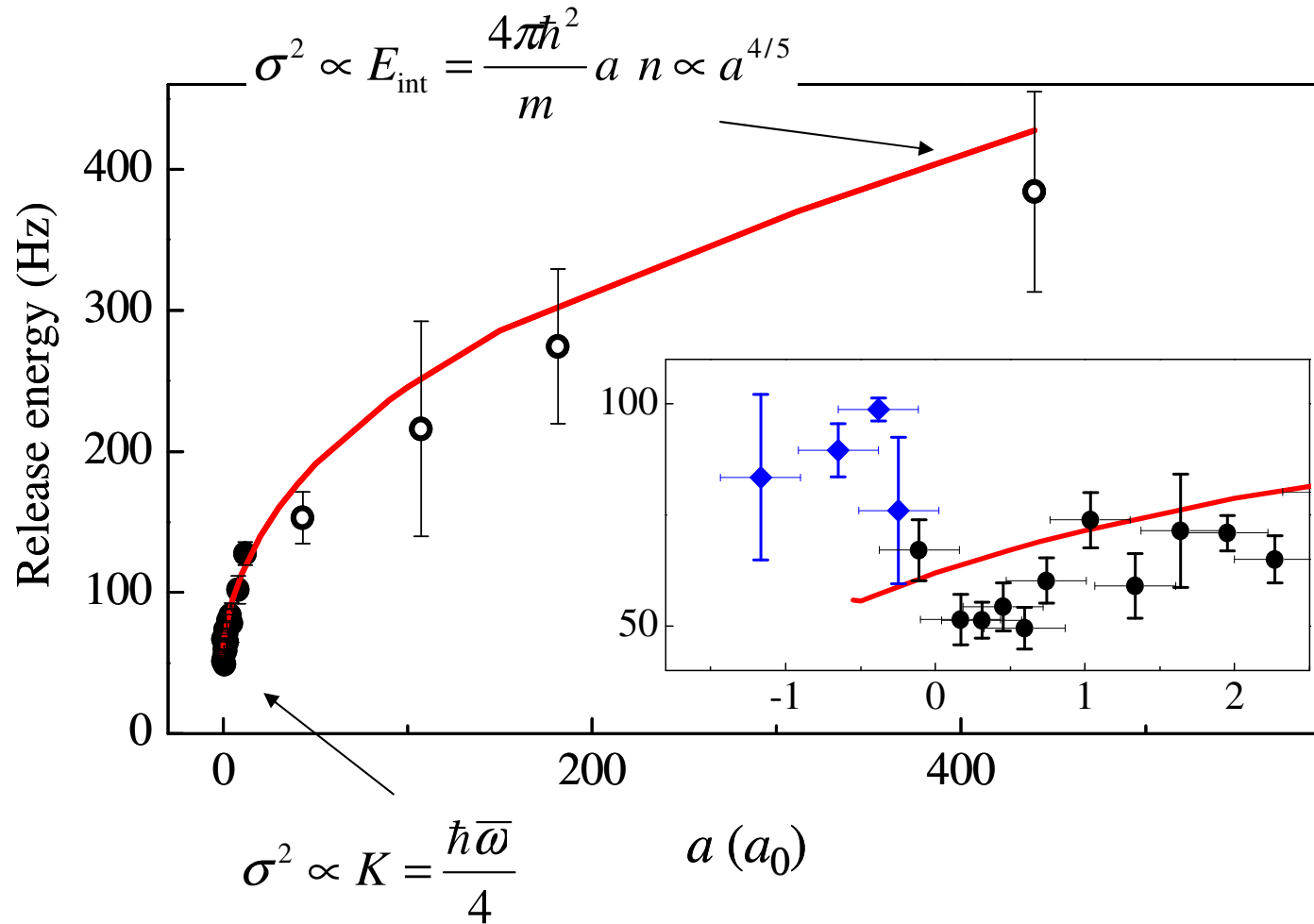
Recent review: C. Chin et al, arXiv:0812.1496

# A Feshbach resonance (potassium-39)



Free expansion:  $E_{\text{int}} \rightarrow K$

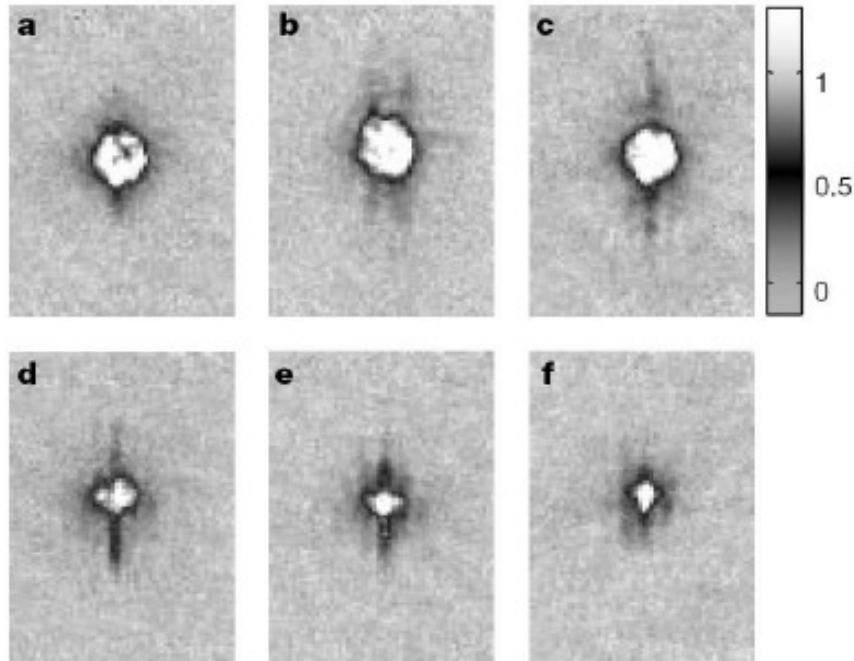
# Interaction energy from free expansion





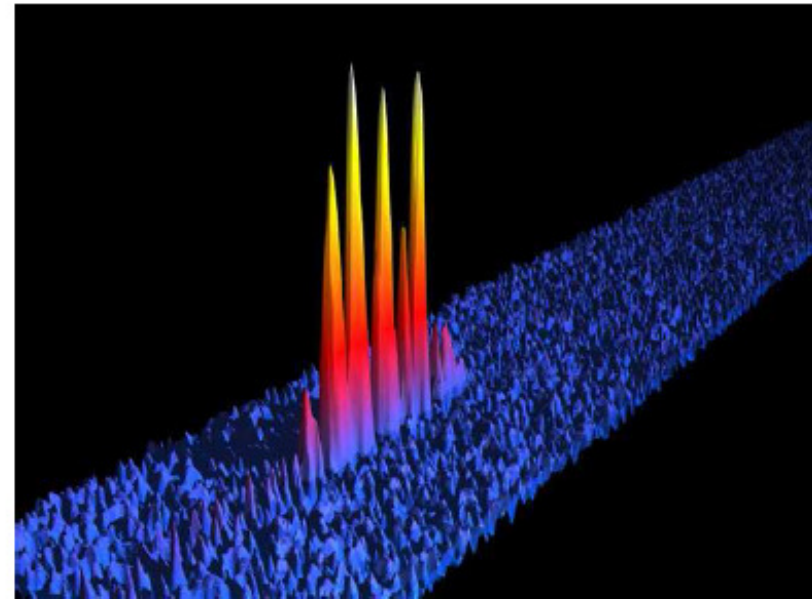
# Attractive interaction

Collapse for  $a < -0.57 \frac{a_{ho}}{N}$



Donley et al., Nature 412, 295 (2001)

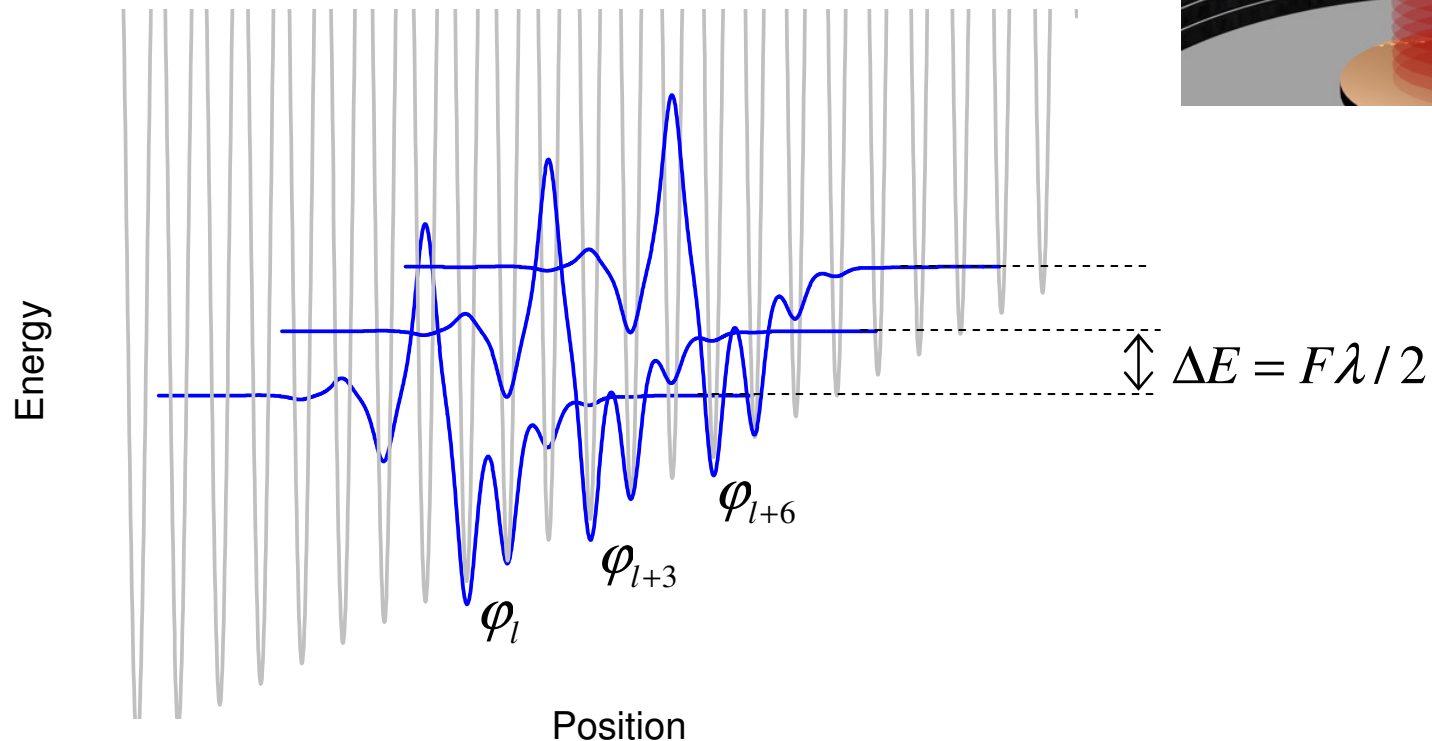
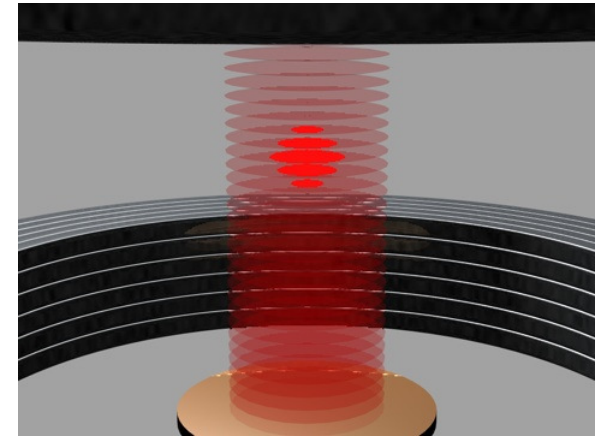
Solitons for intermediate attraction



Strecker et al. Nature 417, 150 (2002); Khaykovich et al. Science 296, 1290 (2002).

# Interaction energy from interferometry

Take a series of equally-spaced energy states...



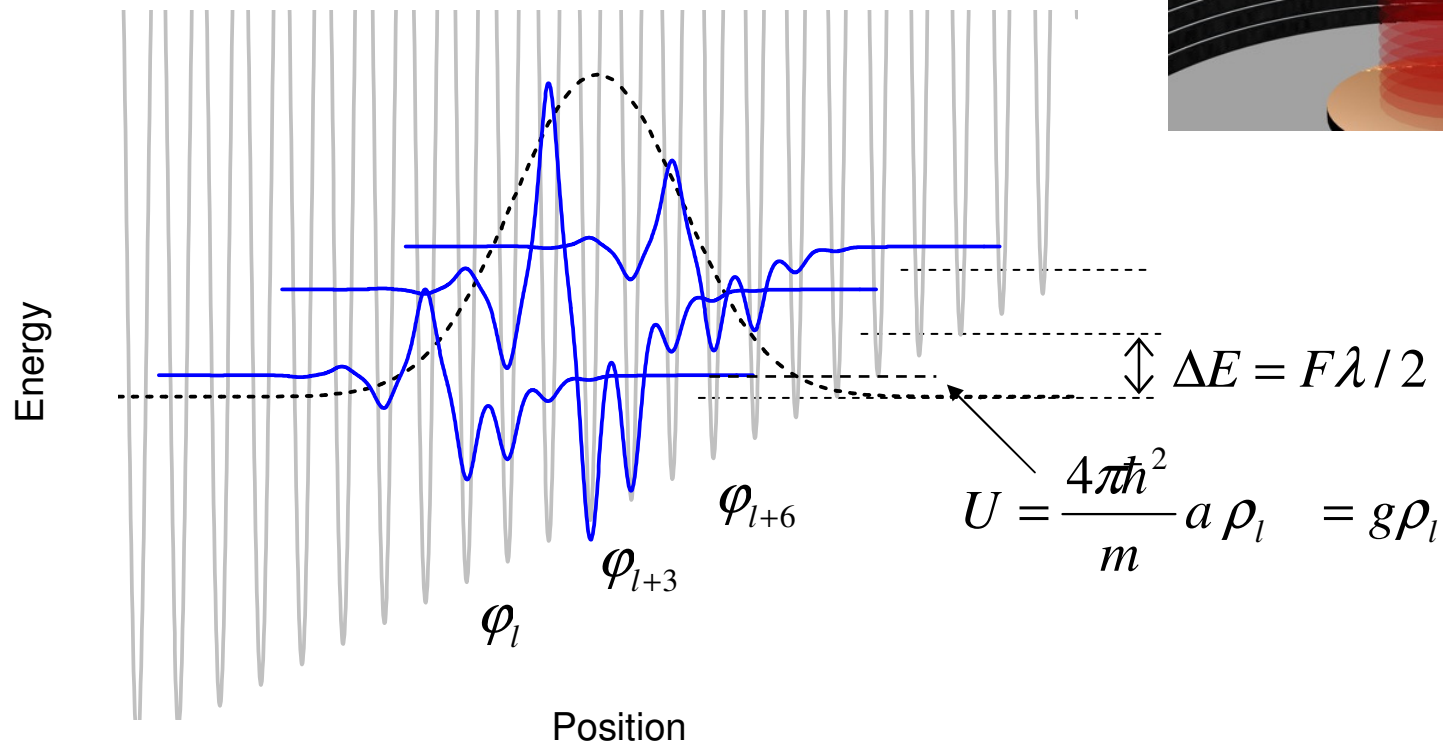
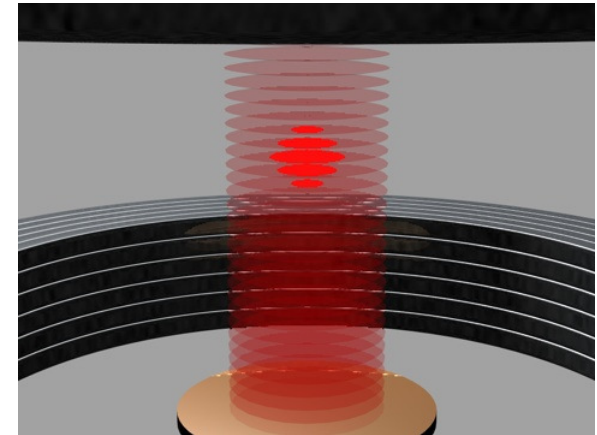
Let them interfere: Bloch oscillations

$$\psi = \sum_l \sqrt{\rho_l} e^{i\vartheta_l} \varphi_l$$

$$\vartheta_l(t) = \frac{Fl\lambda t}{2\hbar}$$

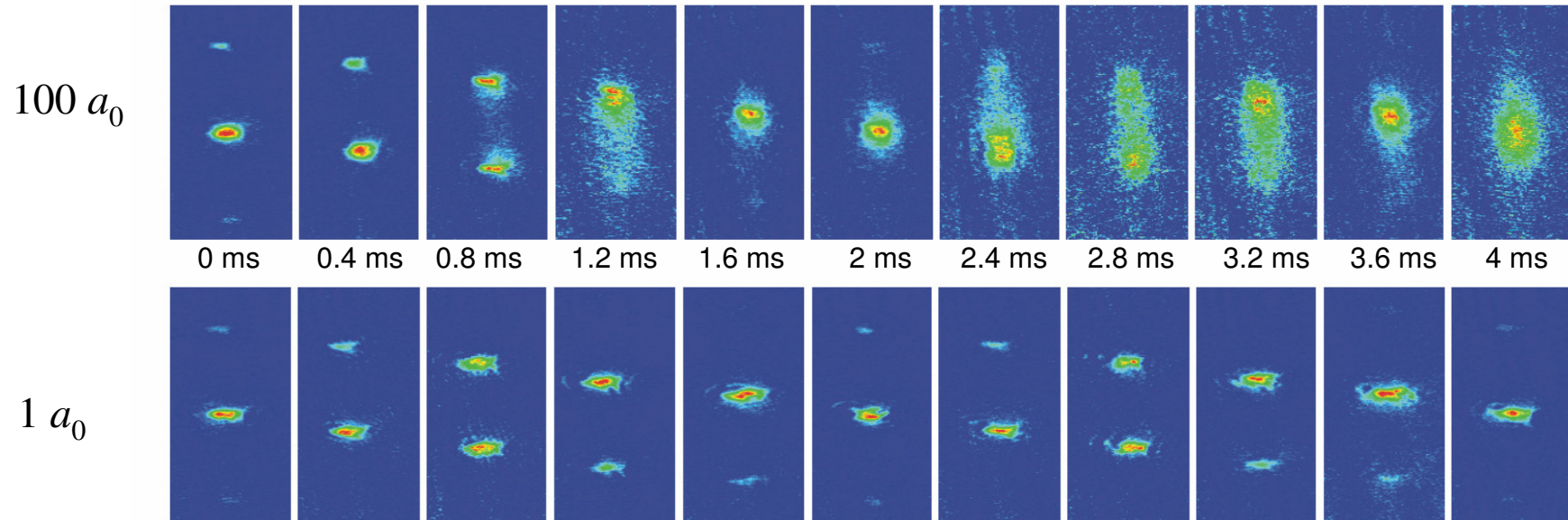
# Interaction energy from interferometry

Add a repulsive interaction ...



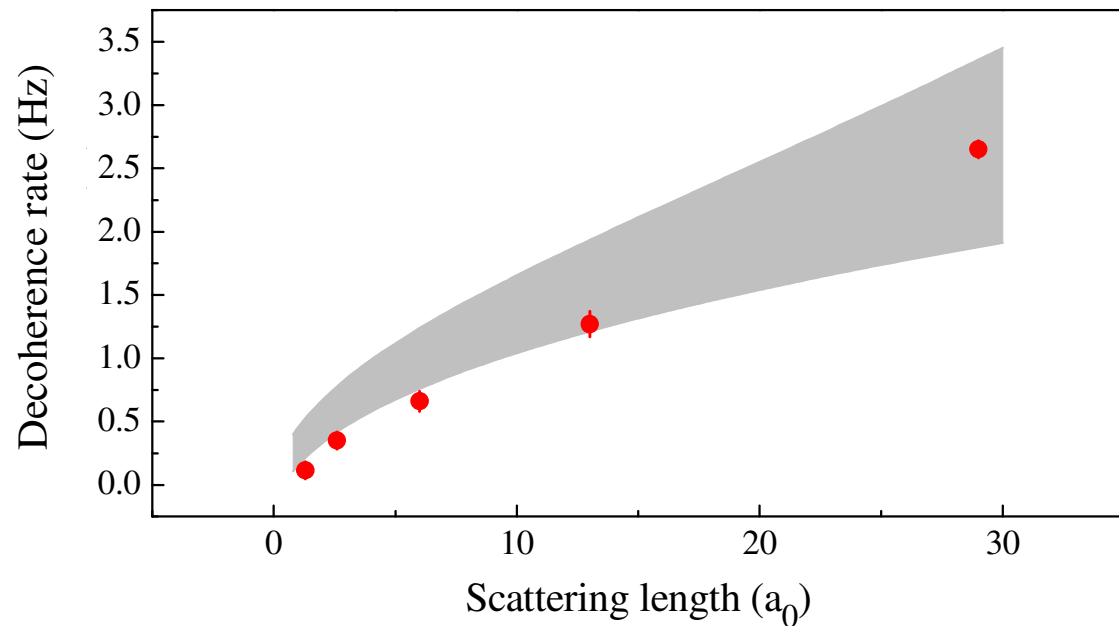
Inhomogeneous broadening:  $\psi = \sum_l \sqrt{\rho_l} e^{i\vartheta_l} \varphi_l$        $\vartheta_l(t) = (Fl\lambda/2 + g\rho_l)t / \hbar$

# Interaction energy from interferometry



To what level can  $E_{int}$  be cancelled?

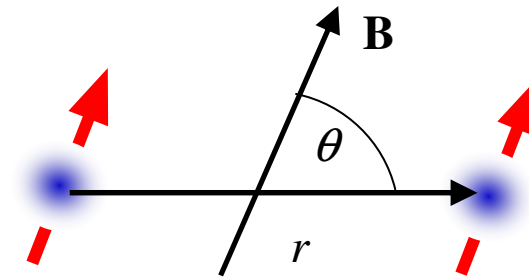
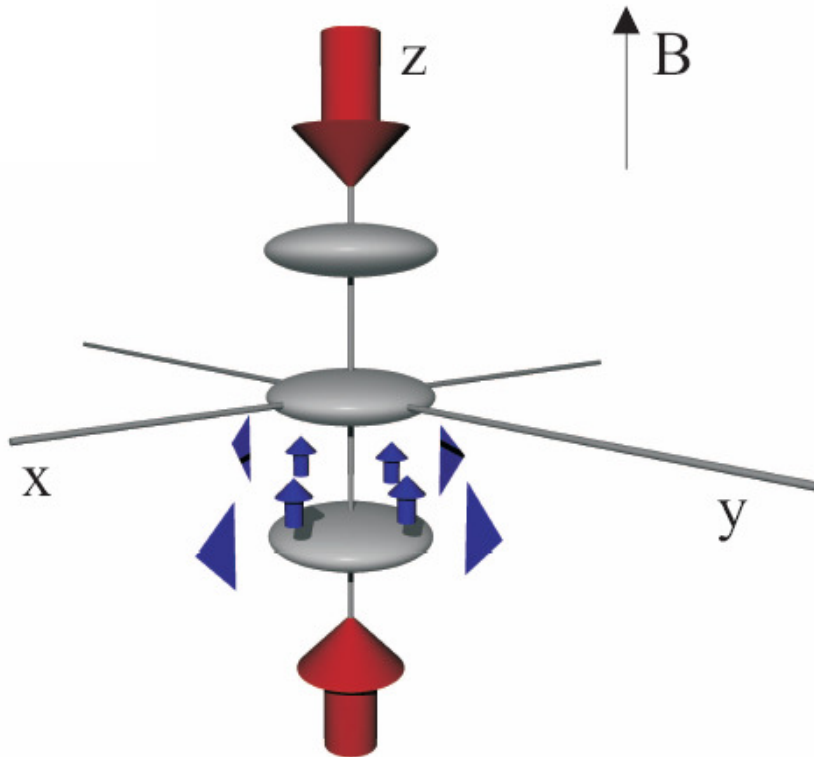
Dipole-dipole interaction is the next order



# Magnetic dipolar interaction

Anisotropic and long-range

$$U_d = \frac{\mu_0 \mu^2}{4\pi r^3} (1 - 3 \cos^2 \vartheta)$$

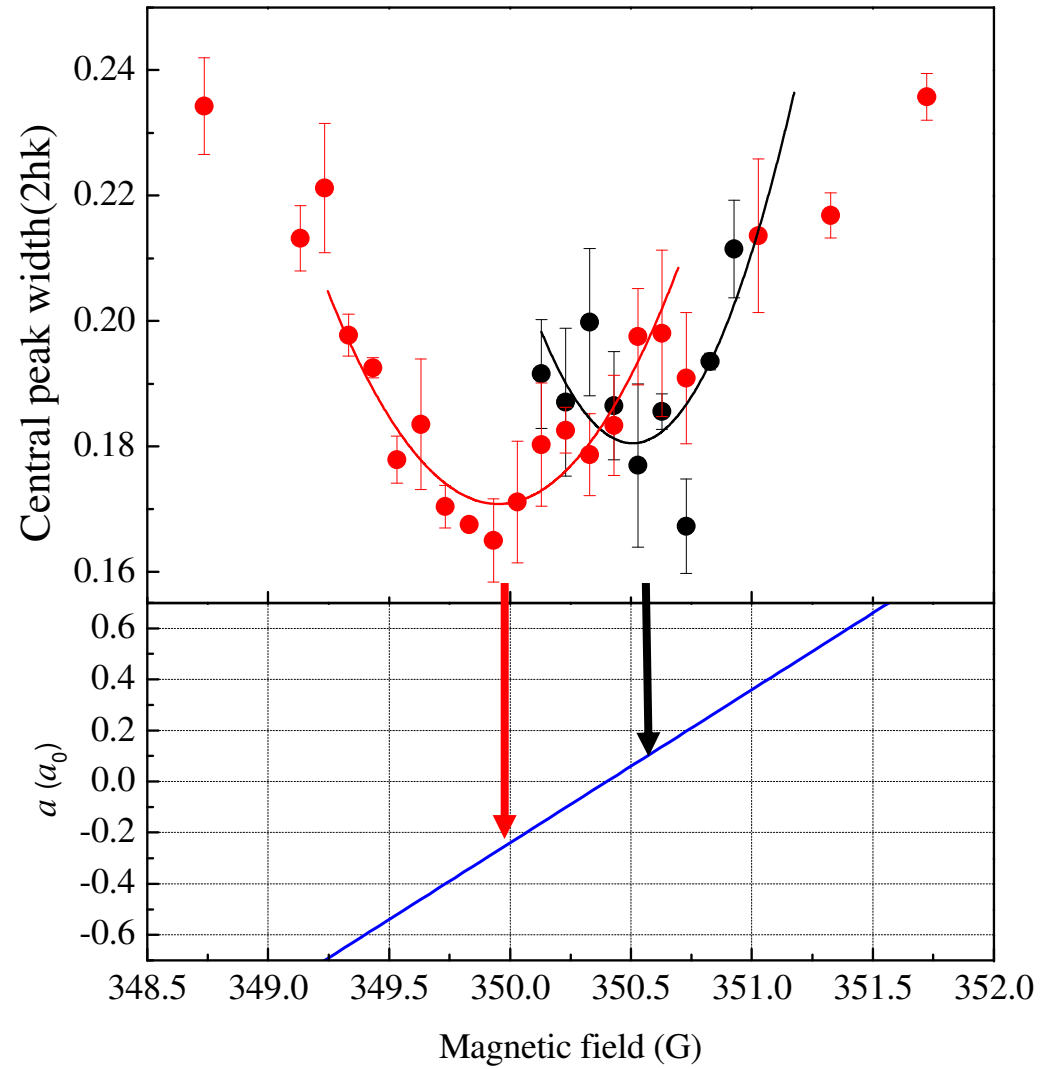
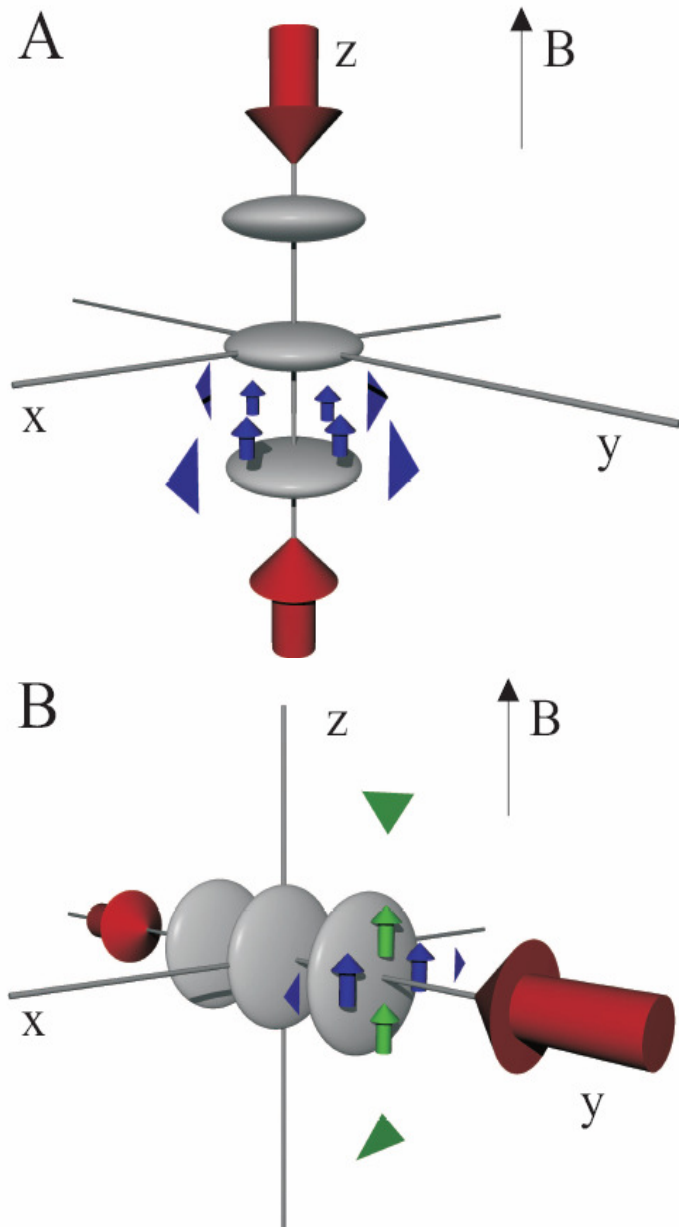


Repulsive dipolar interaction

$$\mu = 0.95 \mu_B \quad r = 0.2 \mu\text{m}$$

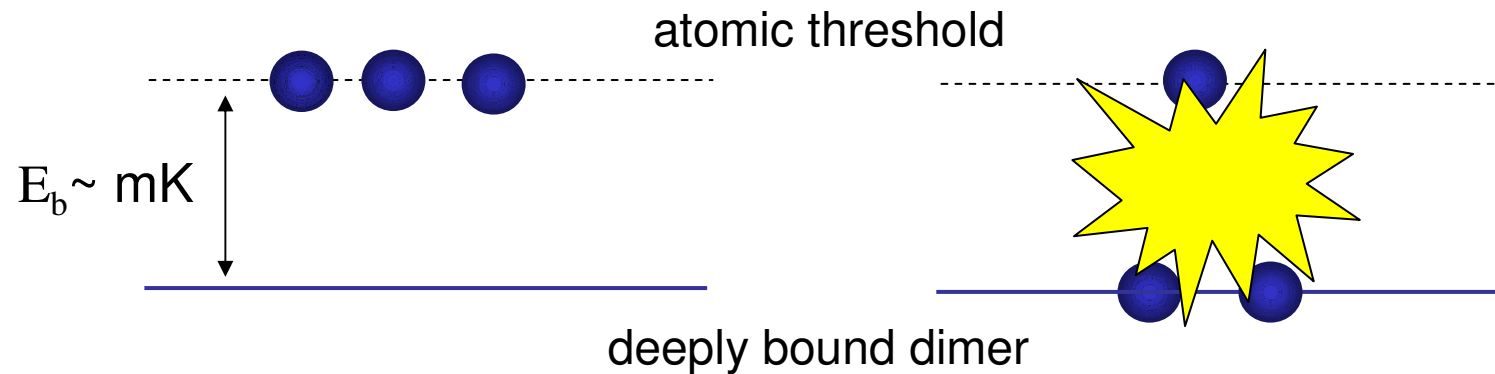
$$U_d \sim 0.5 a_0 \sim 10 \text{ Hz}$$

# Magnetic dipolar interaction: experiment



# Three-body losses

Three atoms can recombine into a molecule+atom: three-body loss



$$\dot{n}(t) = -K_3 n^3(t)$$

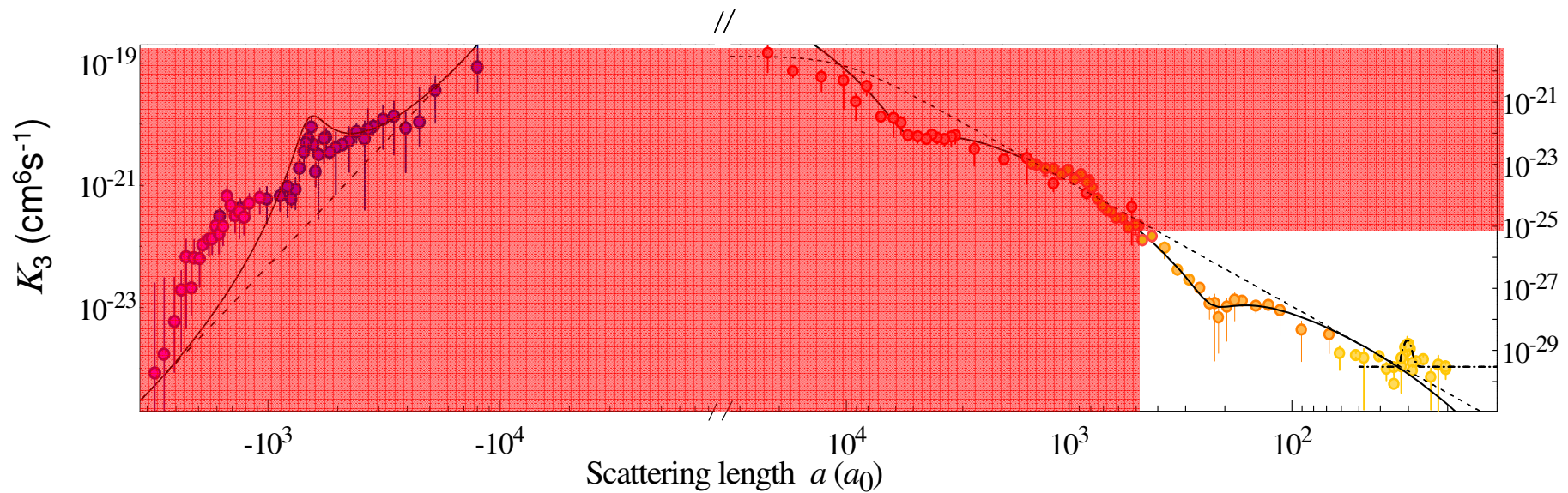
Large  $a$  enhance the recombination at short internuclear distances

$$K_3 \propto \frac{\hbar a^4}{m}$$

The useful range of  $a$  values is limited

# Three-body losses

The picture can be much more complex, due to Efimov trimer states



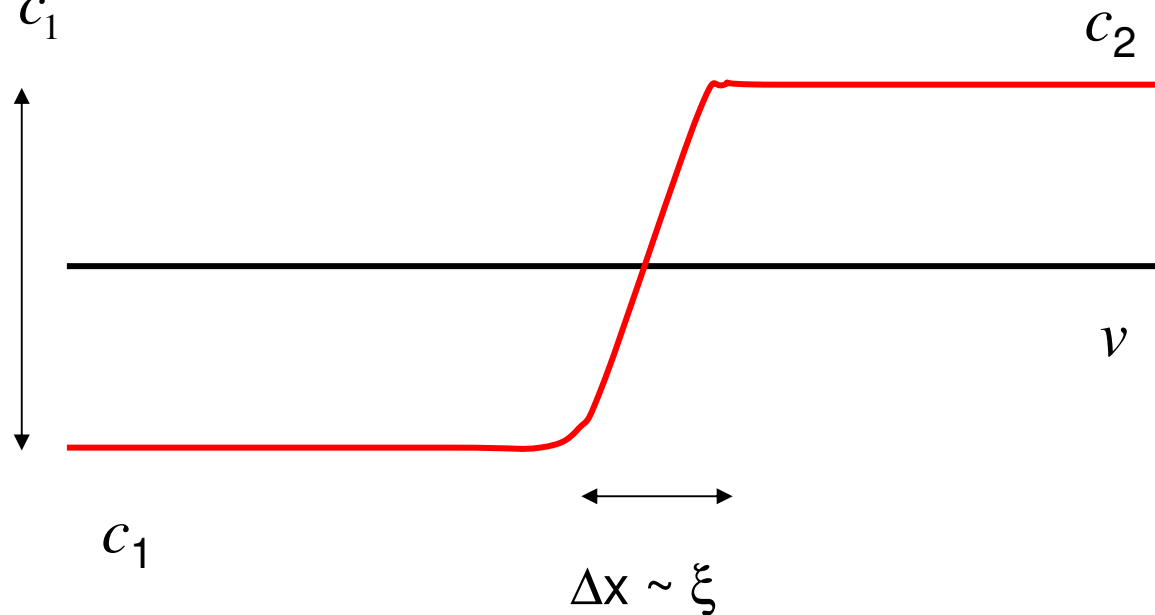
$$n \sim 10^{13} \text{cm}^{-3} \Rightarrow \text{lifetime} < 100 \text{ ms for } K_3 < 3 \times 10^{-26}$$



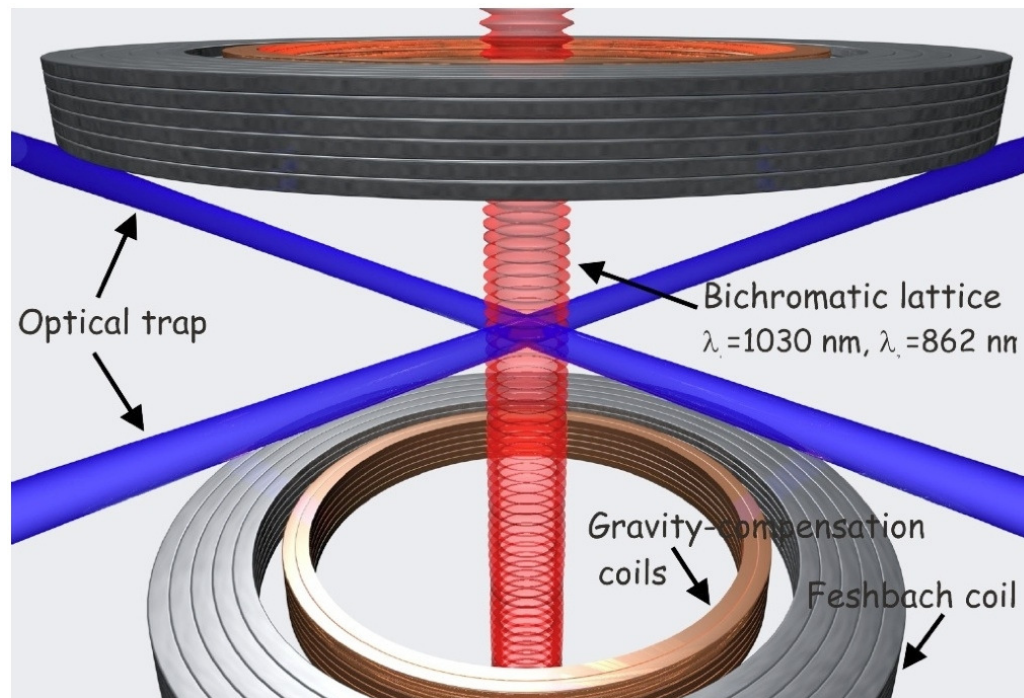
# Feshbach resonances and Hawking radiation

Can one employ Feshbach resonances to make an interaction step in the prescribed way?

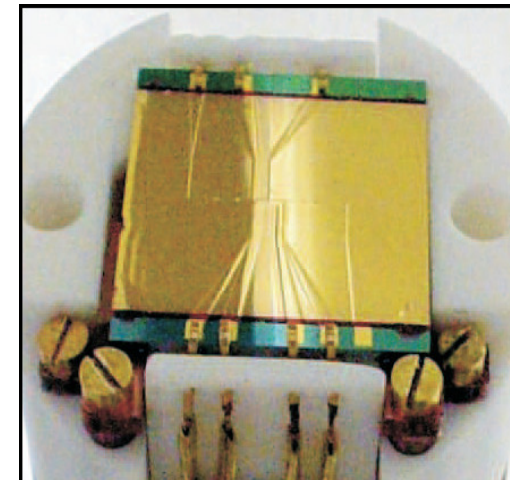
$$D = \frac{c_2 - c_1}{c_1} < 0.5 \quad \Rightarrow \quad a \rightarrow 2a$$



# Feshbach resonances and Hawking radiation



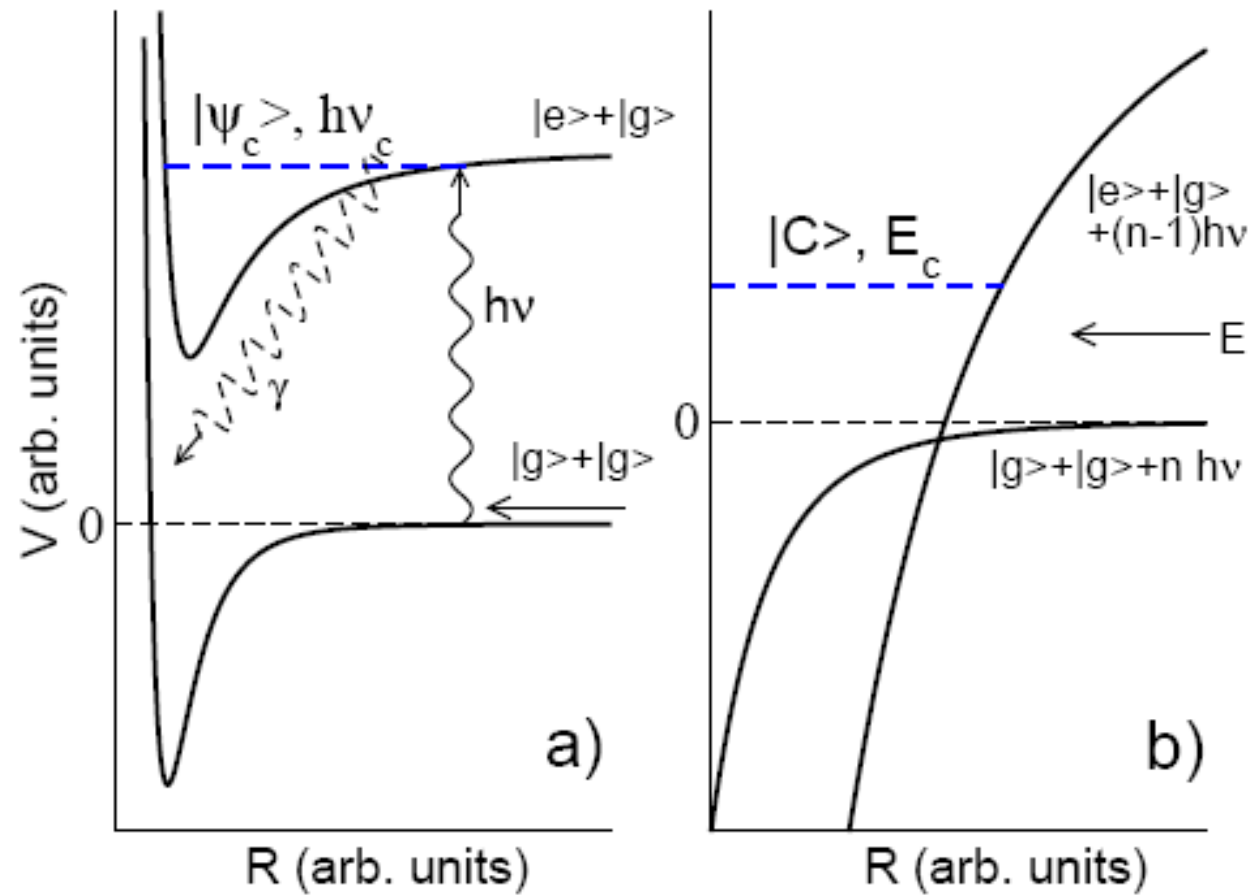
The magnetic coils are typically macroscopic



Easier for atom chips, but still hard

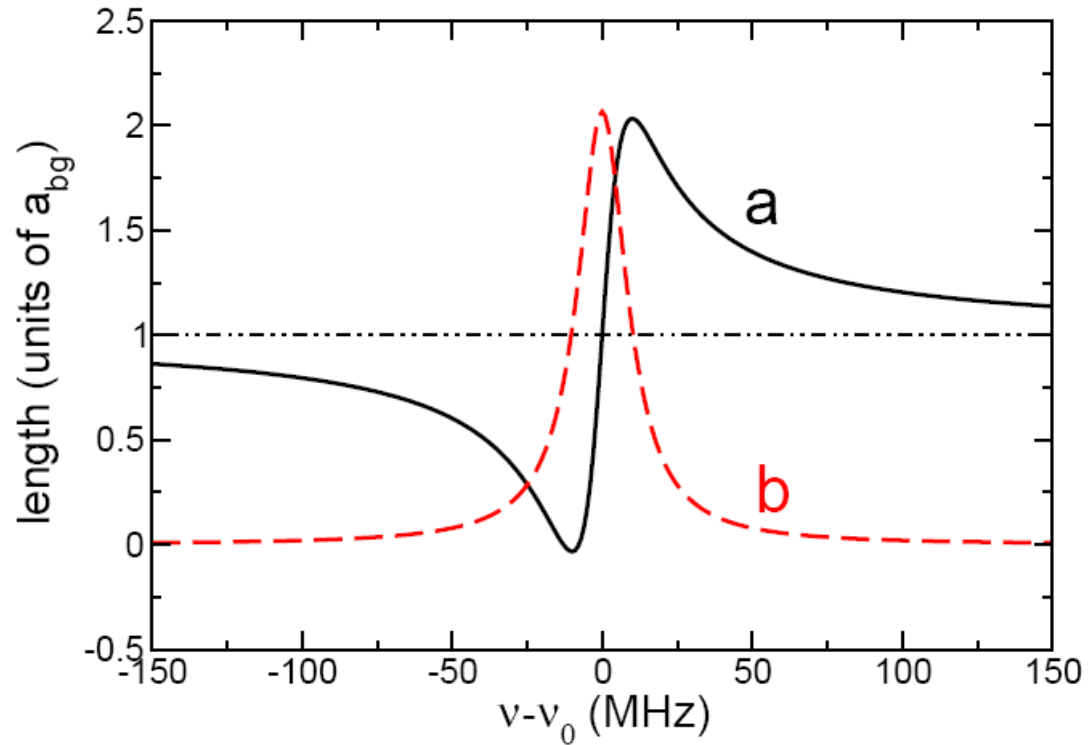
# Optical Feshbach resonance

Laser beams can be focused down to a few  $\mu\text{m}$ , and scanned.



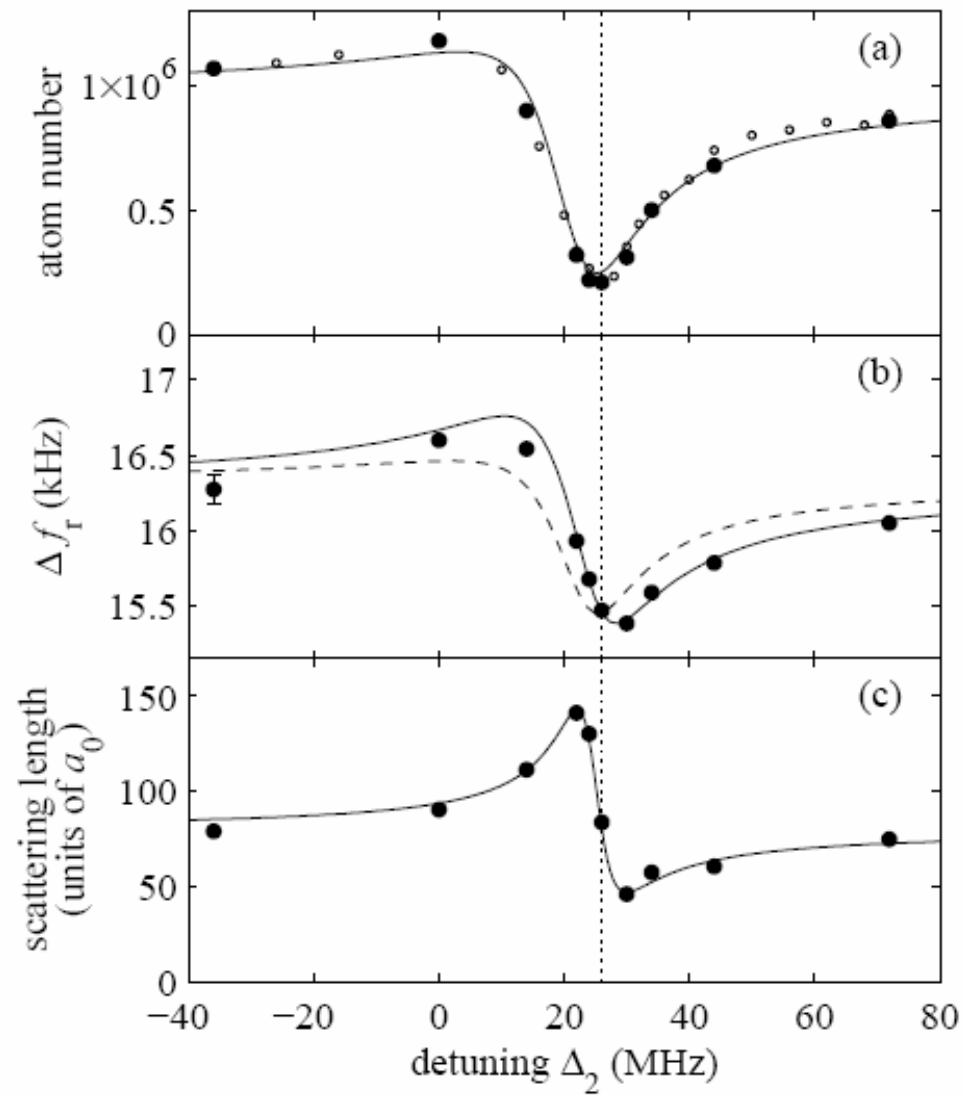
Fedichev et al. Phys. Rev. Lett. 77, 2913 (1996).

# Optical Feshbach resonance



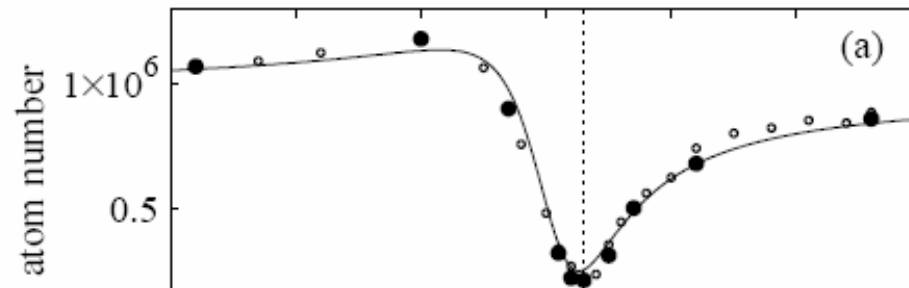
Unfortunately, there is spontaneous emission:  $a \rightarrow a - ib$

# Optical Feshbach resonance



Theis et al, Phys. Rev. Lett. 93, 123001 (2005)

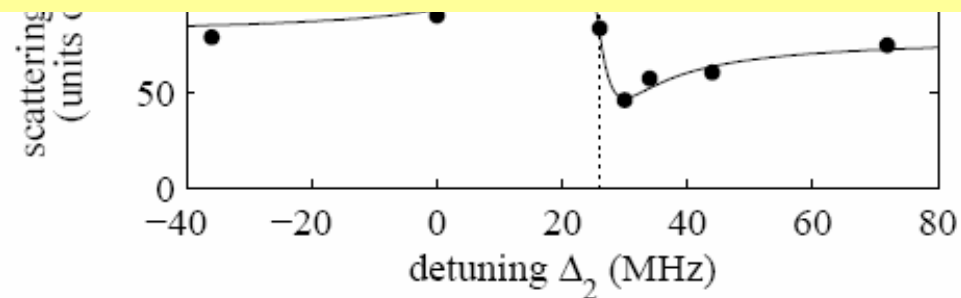
# Optical Feshbach resonance



Alkaline-earth atoms are more promising, because of the long lifetime of selected excited states and good Franck-Condon factors

Yb: 30 nm change of  $a$

Enomoto et al. Phys. Rev. Lett. 101, 203201 (2008)



Theis et al, Phys. Rev. Lett. 93, 123001 (2005)